

NR effective theory for DM direct detection

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Outline

- 🌐 **Overview of dark matter detection**
- 🌐 **Motivation**
- 🌐 **NR effective theory of DM direct detection**
 - Scales and power counting rules
 - Examples of matching
- 🌐 **Recoil spectra**
- 🌐 **Constraints from current direct detection experiments**
- 🌐 **Conclusion**

Overview

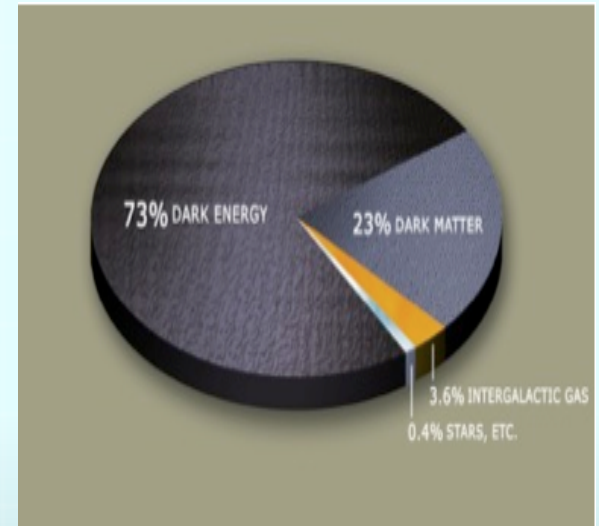
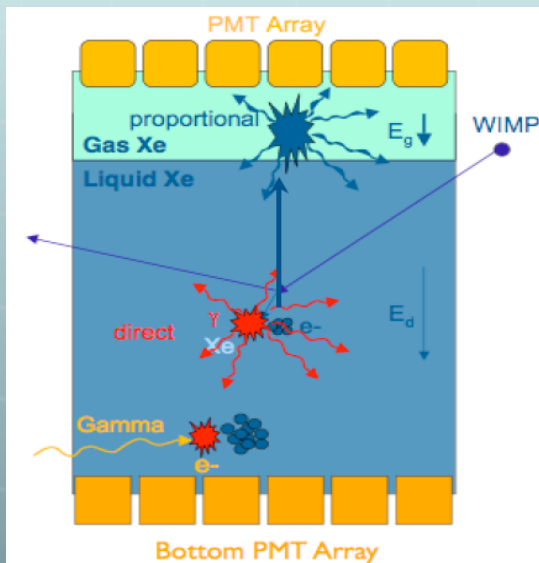
Measurement from CMB + supernovae
+LSS indicates **23% of our universe is composed of DM;**

Three ways to detect DM:

Direct detection

Collider production

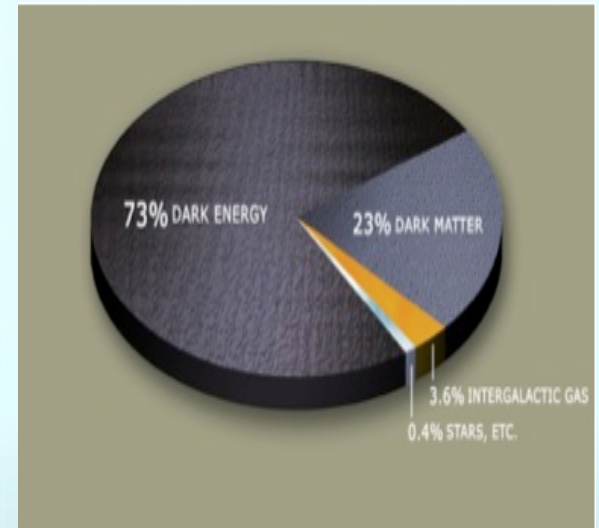
Indirect detection



Overview

Measurement from CMB + supernovae
+LSS indicates **23% of our universe is composed of DM;**

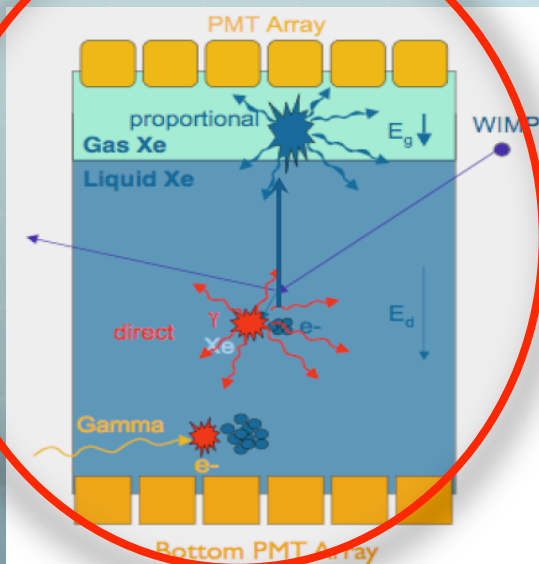
Three ways to detect DM:



Direct detection

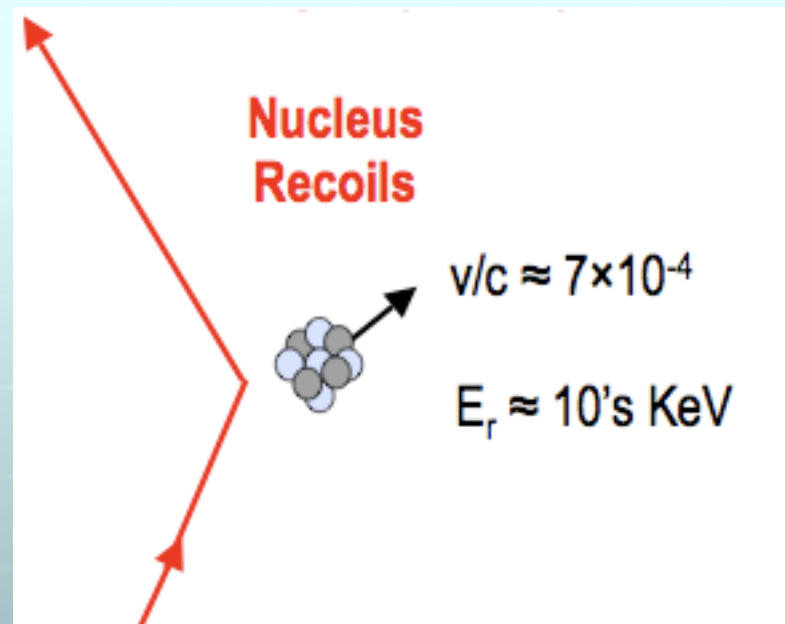
Collider production

Indirect detection



Direct detection

Direct detection looks for signals from DM scattering off nucleus in the underground detector.



Dark matter



Direct detection kinematic regime

$$E_{kin} \sim \mathcal{O}(10 \text{ keV})$$

$$E_R \sim \mathcal{O}(1 - 10 \text{ keV})$$

The scattering against the whole nucleus is described by non-relativistic quantum mechanics.

Current direct detection status

Direct detection rate

$$\frac{dR}{dE_R} = N_T \frac{\rho_\chi}{m_\chi} \int_{v_{min}}^{v_E} d^3v v f(v, v_E) \frac{d\sigma}{dE_R}$$

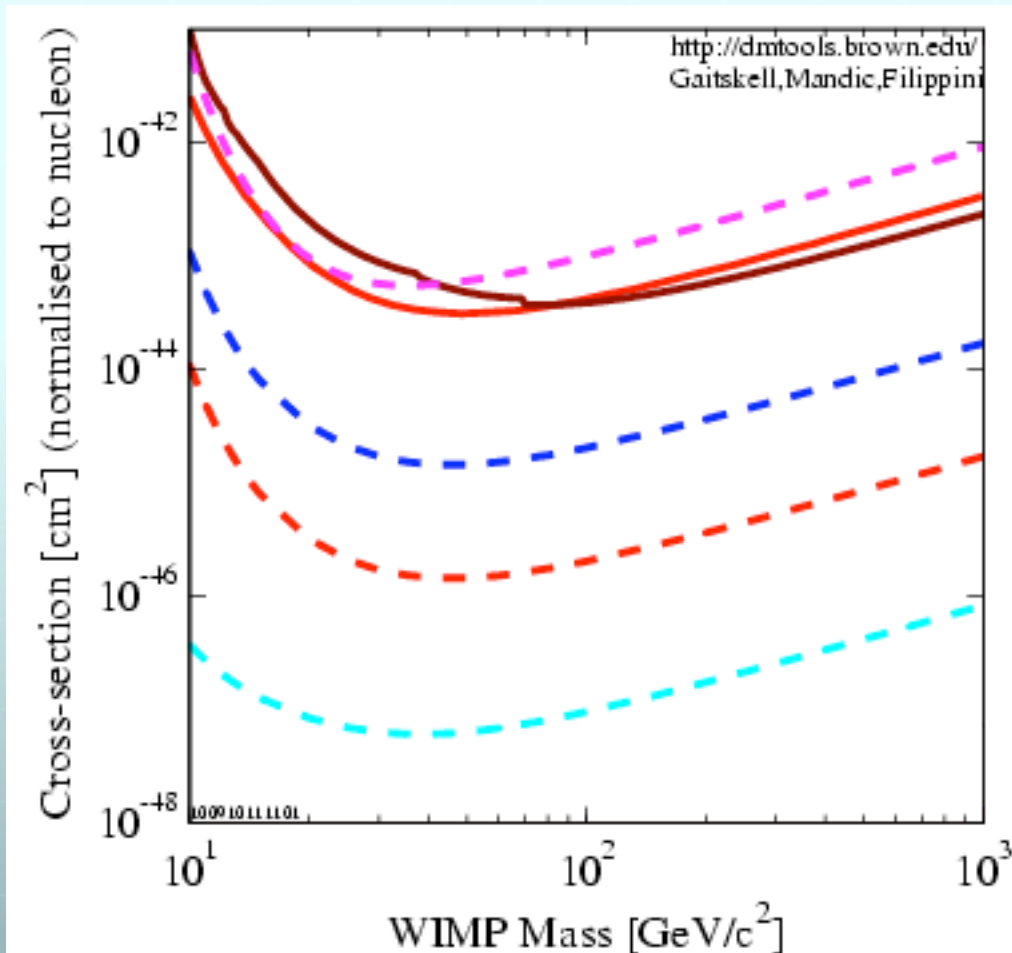
Traditionally, signals are assumed to come from contact interactions independent of momentum transfer. For instance, for SI scattering,

$$\frac{d\sigma}{dE_R} = \frac{m_N}{2v^2} \frac{\sigma_n}{\mu_n^2} A^2 F^2(E_R)$$

Experimental results are presented as bounds on σ_n as a function of DM mass

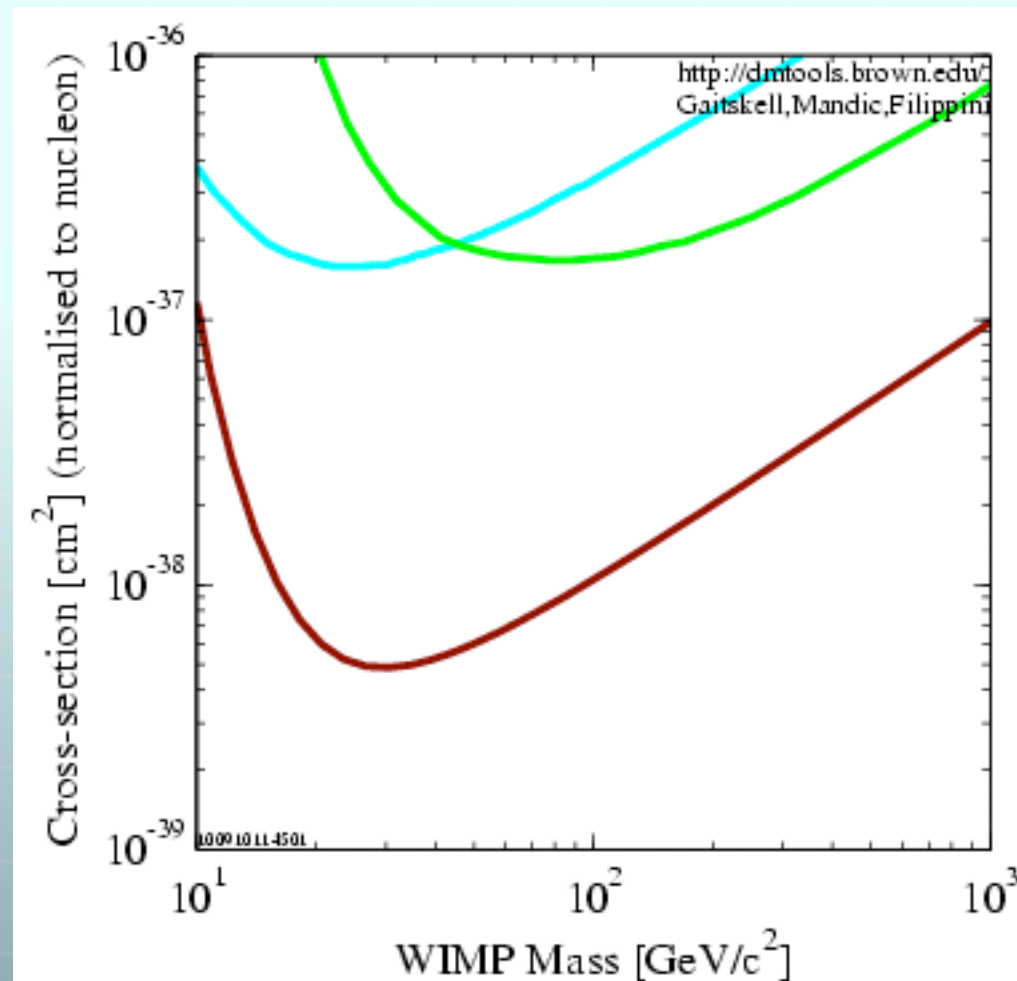


Spin-independent elastic scattering



- DATA listed top to bottom on plot
- XENON10 2007, measured σ_{eff} from Xe cube
- CDMS: Soudan 2004-2009 Ge
- XENON100 2010 (161 kg-d)
- XENON100 projected sensitivity: 6000 kg-d, 5-30 keV, 45% eff.
- XENON100 upgrade projected sensitivity: 60,000 kg-d, 5-30 keV, 45% eff.
- XENON 1T projected sensitivity: 3 ton-yr, 2-30 keV, 45% eff.

Continued: spin-dependent scattering



DATA listed top to bottom on plot
KIMS 2007 - 3-409 kg-days CsI SD-proton
PICASSO SD-proton (2009)
XENON10 SD-neutron
100910114501

Summary

 The current bounds on xsec per nucleon are

SI scattering: $(10^{-43} - 10^{-44})cm^2$

SD scattering: $(10^{-37} - 10^{-38})cm^2$

Motivation

- 🌐 **How to extract more information about the underlying dark dynamics from the experimental data?**
- 🌐 **The simple assumption of contact interaction between DM with nucleons overlooks direct detection's sensitivity to more general DM scenarios.**

E.g: Momentum-dependent DM (Chang, Pierce and Weiner; Feldstein, Fitzpatrick and Katz 2009);

Dark electromagnetic moments (Chang, Weiner and Yavin; Barger, Keung and Marfatia; Fitzpatrick and Zurek; Banks, Fortin and Thomas 2010)

Top-down view: (from theory to direct detection experiment)

UV complete DM Models predict relic abundance, direct detection signals in terms of specific model parameters



Effective field theory operators (Kurylov and Kamionkowski 2003;

NR limit

Agrawal, Chacko, Kilic and Mishra 2010)



Model independent analysis

Non-relativistic quantum mechanics theory

- 🌐 **Different models/effective ops lead to the same NR interaction for direct detection.**

E.g: Higgs exchange: $\bar{\chi}\chi\bar{q}q$

Z exchange: $\bar{\chi}\gamma^\mu\chi\bar{q}\gamma_\mu q$

gives the contact interaction in the NR limit.

- 🌐 **Measured recoiling rate directly bounds the coupling of contact interaction.**

❑ **From direct detection experiments to theories:**

Direct detection



Non-relativistic effective potential



Effective field theory operators

UV complete DM Models

NR effective potential

Scales and simple power counting

Transferred momentum: $|q| \sim 100 \text{ MeV}$

Nucleus mass: $m_N \sim 10 - 100 \text{ GeV}$

DM mass: $m_\chi \sim 100 \text{ GeV} - 1 \text{ TeV}$

Mediator mass: m_0 unfixed

Other scales: e.g. DM-mediator interaction arises at nonrenormalizable level

DM with electric dipole moment

$$\frac{\bar{\chi} \sigma_{\mu\nu} \gamma_5 \chi F^{\mu\nu}}{\Lambda}$$

Continued

Expansion parameters:

$$v \sim 10^{-3}$$

$$\frac{q}{m_N} \sim \frac{q}{m_\chi} \sim 10^{-3}$$

$\frac{q}{m_0}$, $\frac{q}{\Lambda}$ **unfixed, can be as large as 0.1**

 **SI experimental bounds**

$$\sigma \sim 10^{-44} \text{cm}^2$$

 **SD experimental bounds**

$$\sigma \sim 10^{-38} \text{cm}^2$$

compared to a typical weak process xsec

$$\sigma_W \sim 10^{-36} \text{cm}^2$$

|q| suppressed operator can still be relevant if they are the leading operator for direct detection.

 **Consider two limits of mediator masses m_0 :**

a. $m_0 \gg |q|$ Contact interaction

b. $m_0 \ll |q|$ Long-range interaction

Assume all expansion parameters of order 10^{-3}

a. Contact interaction: Operators suppressed by a single $|q|$

b. Long-range interaction: Operators suppressed by $|q|^3$

may still be relevant for direct detection if they are the leading operator.

Effective NR potential

$$V_{\text{eff}} = V_{\text{eff}}^{\text{SI}} + V_{\text{eff}}^{\text{SD}}$$

$$V_{\text{eff}}^{\text{SI}} = h_1 \delta^3(\vec{r}) - h_2 \vec{s}_\chi \cdot \vec{\nabla} \delta^3(\vec{r})$$

$$+ l_1 \frac{1}{4\pi r} + l_2 \frac{\vec{s}_\chi \cdot \vec{r}}{4\pi r^3},$$

$$V_{\text{eff}}^{\text{SD}} = h'_1 \vec{s}_\chi \cdot \vec{s}_N \delta^3(\vec{r}) - h'_2 \vec{s}_N \cdot \vec{\nabla} \delta^3(\vec{r})$$

$$+ l'_1 \frac{\vec{s}_\chi \cdot \vec{s}_N}{4\pi r} + l'_2 \frac{\vec{s}_N \cdot \vec{r}}{4\pi r^3},$$

contact interaction

long range interaction

- ❑ In the Born approximation, the matrix element of the scattering is

$$\mathcal{M}(\vec{q}, \vec{v}) = - \int d^3\vec{r} e^{i\vec{q}\cdot\vec{r}} V_{eff}(\vec{r}, \vec{v})$$

- ❑ For the numerical studies, we only consider elastic scattering;
- ❑ Only list static potential, or, v-independent potential;
v-dependent potential produces nearly identical recoil spectrum to the static one for the elastic scattering;
- ❑ Only list potential leads to interaction suppressed by a single $|q|$;
- ❑ The coefficients are dimensionful; direct detection bounds on combinations of couplings and scales.
- ❑ We factor out the nuclear form factor and suppress the spin indices.

Example of matching: Fermionic DM(SI)

 **SI NR operator complete set:** up to a scalar function $f(q^2, v^2)$

In general, four building blocks to construct rotational invariants

$$\vec{q}, \vec{v}, \vec{s}_\chi, \vec{s}_N$$

$$\vec{r}, \vec{v}, \vec{s}_\chi, \vec{s}_N$$

momentum space (w/o mediator)

$$\begin{matrix} P & C \\ \mathcal{O}_1^{(++)} & = 1 \end{matrix}$$

$$\mathcal{O}_2^{(-+)} = i\vec{s}_\chi \cdot \vec{q}$$

$$\mathcal{O}_3^{(--)} = \vec{s}_\chi \cdot \vec{P}$$

$$\mathcal{O}_4^{(++)} = i\vec{s}_\chi \cdot (\vec{P} \times \vec{q})$$

$$(\vec{P} = \mu_N \vec{v} + \vec{q}/2)$$

position space

$$\mathcal{O}_1 = \delta^3(\vec{r}), \quad \frac{1}{4\pi r}$$


$$\mathcal{O}_2 = \vec{s}_\chi \cdot \vec{\nabla} \delta^3(\vec{r}), \quad \frac{\vec{s}_\chi \cdot \vec{r}}{4\pi r^3}$$

$$\mathcal{O}_3 = \vec{s}_\chi \cdot \vec{P} \delta^3(\vec{r}), \quad \frac{\vec{s}_\chi \cdot \vec{P}}{4\pi r}$$

$$\mathcal{O}_4 = (\vec{s}_\chi \times \vec{v}) \cdot \delta^3(\vec{r}), \quad \frac{(\vec{s}_\chi \times \vec{v}) \cdot \vec{r}}{4\pi r^3}$$



Mapping of effective field theory operators to the NR potential

$$\begin{aligned} \mathcal{O}_1^{(++)} &= 1 && \bar{\chi}\chi\bar{q}q, \quad \bar{\chi}\gamma^\mu\chi\bar{q}\gamma_\mu q \\ \mathcal{O}_2^{(-+)} &= i\vec{s}_\chi \cdot \vec{q} && \bar{\chi}\gamma^5\chi\bar{q}q \quad \bar{\chi}\sigma^{\mu\nu}\gamma^5 D_\mu\chi\bar{q}\gamma_\nu q \\ \mathcal{O}_3^{(--)} &= \vec{s}_\chi \cdot \vec{P} && \bar{\chi}\gamma^5\gamma^\mu\chi\bar{q}\gamma_\mu q \\ \mathcal{O}_4^{(++)} &= i\vec{s}_\chi \cdot (\vec{P} \times \vec{q}) && \bar{\chi}\sigma^{\mu\nu} D_\mu\chi\bar{q}\gamma_\nu q \end{aligned}$$




Examples of simple theories for each potential

$$\delta^3(\vec{r})$$

Higgs exchange

$$\vec{s}_\chi \cdot \vec{\nabla} \delta^3(\vec{r})$$

DM dark EDM off nucleus charge

$$\frac{1}{r}$$

exchange of a light boson

$$\frac{\vec{s}_\chi \cdot \vec{r}}{r^3}$$

DM EDM off nucleus charge

Recoil spectrum

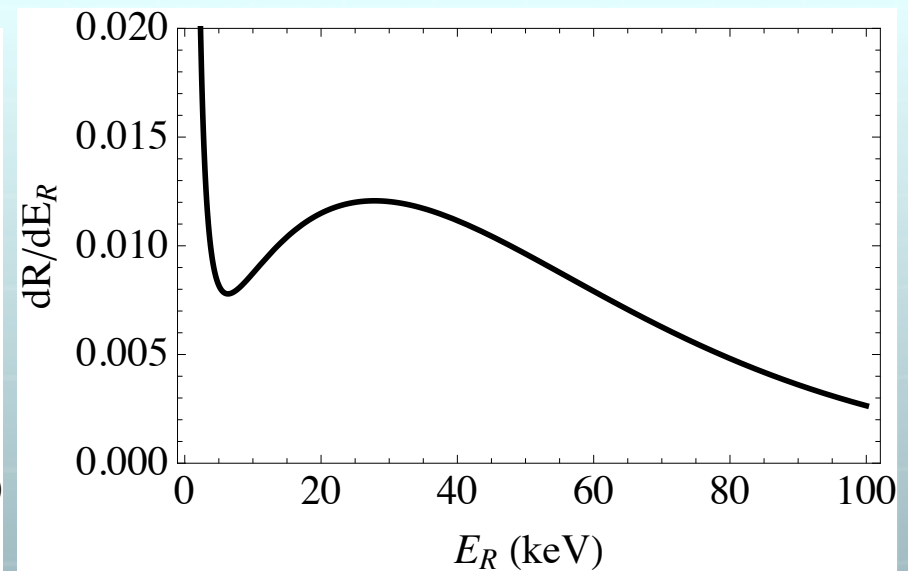
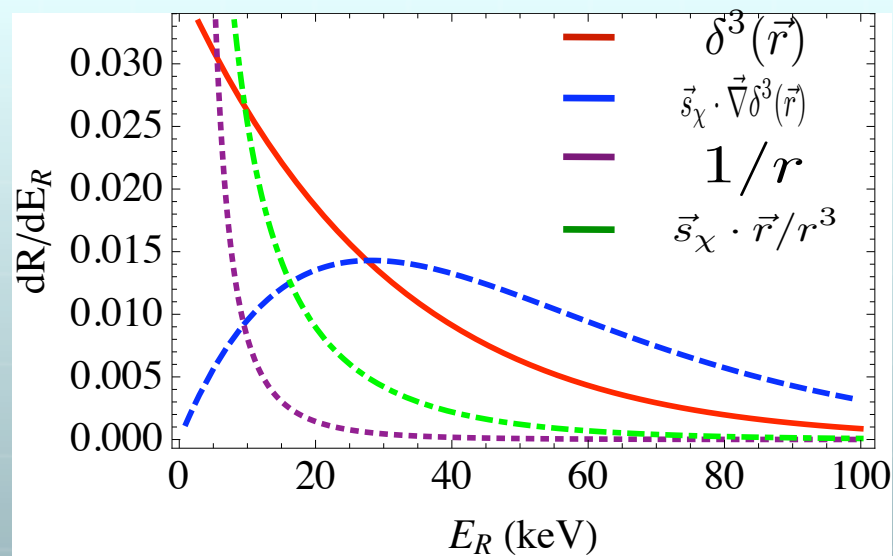
The NR theory highlights the possibility of having qualitatively different recoil energy spectrum.

SI NR operators	SD NR operators	E_R
$\delta^3(\vec{r})$	$\vec{s}_\chi \cdot \vec{s}_N \delta^3(\vec{r})$	1
$\vec{s}_\chi \cdot \vec{\nabla} \delta^3(\vec{r})$	$\vec{s}_N \cdot \vec{\nabla} \delta^3(\vec{r})$	E_R
$\frac{1}{4\pi r}$	$\frac{\vec{s}_\chi \cdot \vec{s}_N}{4\pi r}$	E_R^{-2}
$\frac{\vec{s}_\chi \cdot \vec{r}}{4\pi r^3}$	$\frac{\vec{s}_N \cdot \vec{r}}{4\pi r^3}$	E_R^{-1}

Sample Spectra

🌐 Sample spectrum for Germanium.

Left: spectrum with contribution from one operator;



Right: spectrum with contributions from two operators;

**Constraints from direct detection (CDMS, Xenon10,
Xenon100 for SI direct detection)**

$$\delta^3(\vec{r}) \quad h_1 \lesssim 10^{-8} \text{ GeV}^{-2} = \frac{10^{-4}}{(100 \text{ GeV})^2}$$

$$\vec{s}_\chi \cdot \vec{\nabla} \delta^3(\vec{r}) \quad h_2 \lesssim 10^{-7} \text{ GeV}^{-3} = \frac{10^{-1}}{(100 \text{ GeV})^3}$$

$$\frac{1}{r} \quad l_1 \lesssim 10^{-11}$$

$$\frac{\vec{s}_\chi \cdot \vec{r}}{r^3} \quad l_2 \lesssim 10^{-9} \text{ GeV}^{-1} = \frac{10^{-7}}{(100 \text{ GeV})}$$

🌐 **Example 1** $\delta^3(\vec{r})$ $h_1 \lesssim 10^{-8} \text{ GeV}^{-2} = \frac{10^{-4}}{(100 \text{ GeV})^2}$

Higgs exchange: $y_s \sim 10^{-4}$

Z exchange: $\bar{\chi} \gamma^\mu \chi h^\dagger D_\mu h$ $(v_{EW}/\Lambda)^2 \sim 10^{-4}$

🌐 **Example 2** $\vec{s}_\chi \cdot \vec{r}/r^3$ $l_2 \lesssim 10^{-9} \text{ GeV}^{-1} = \frac{10^{-7}}{(100 \text{ GeV})}$

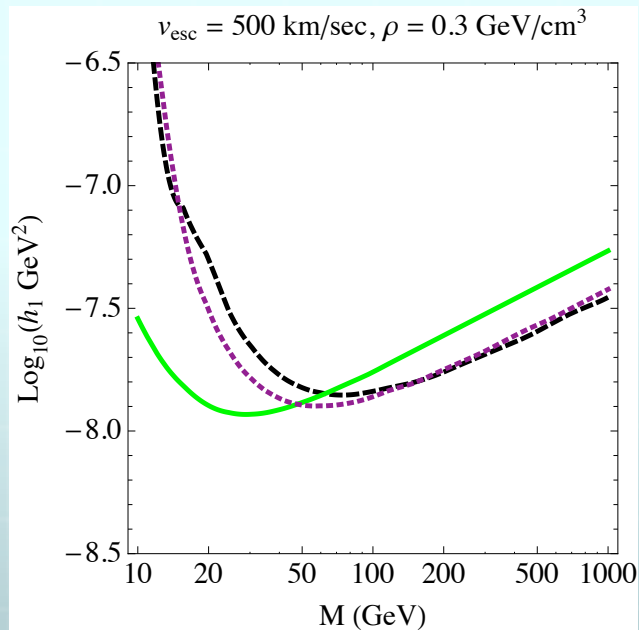
(Pseudo)scalar exchange: $l_2 \sim g/m_\chi$ $g \lesssim 10^{-7}$

DM dipole moment: $l_2 = d$ $d \lesssim 10^{-23} (e \cdot \text{cm})$

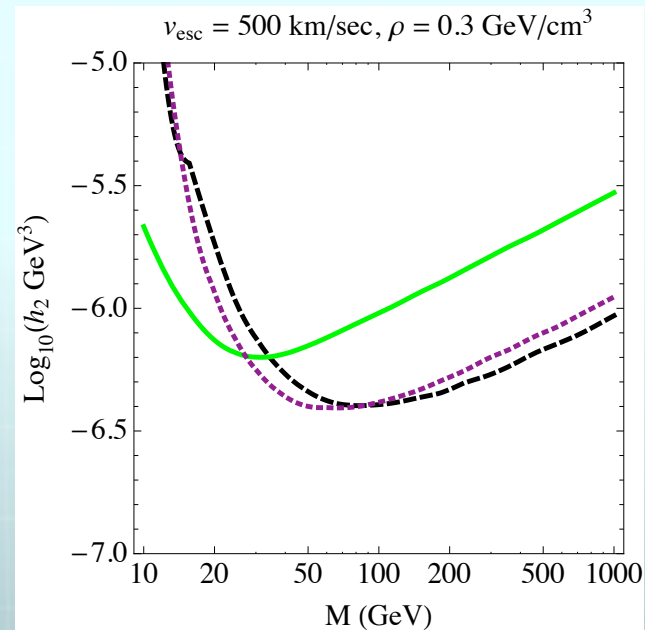


Bound on contact interaction

momentum independent



momentum dependent

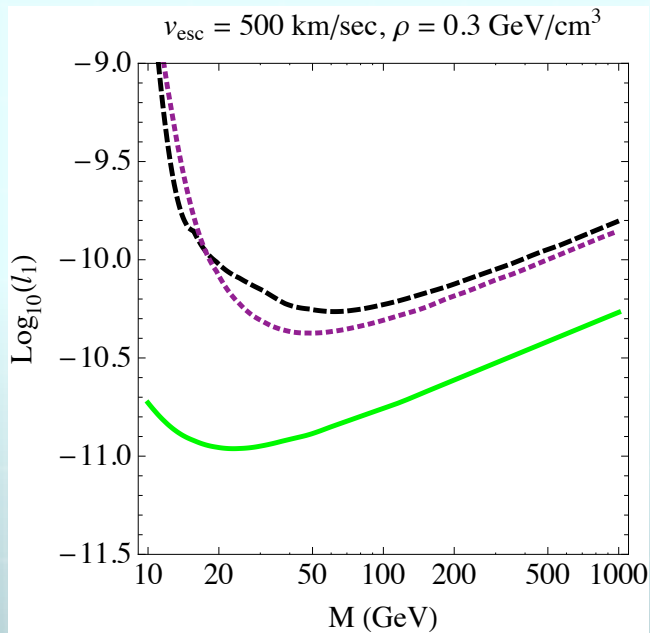


Black: CDMS; Green: Xenon 10; Purple: Xenon100

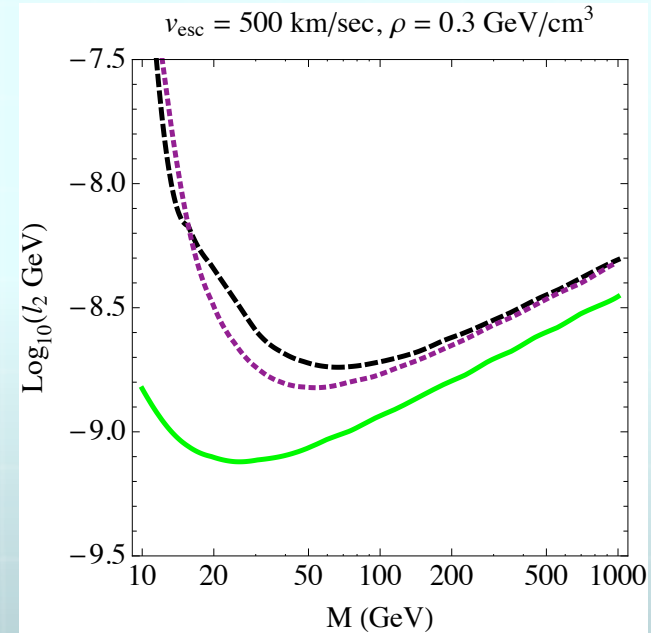


Bound on long range interaction

coloumb potential





dipole potential



Xenon 10 has strongest bound as it has the lowest energy threshold;

Threshold calibration is crucial for not only light DM but also long range interaction mediated by light bosons.

Conclusion

-  We present a model-independent framework based on NR operators to analyze data from direct detection.
-  If near future direct detection sees DM, it will not only shed information on DM mass, overall scattering cross section but also DM interaction from recoiling spectrum.

Thank you!