

# Displaced Leptons and Other Exotic Objects at the LHC

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Evans, Shelton – arXiv:1601.01326

## Status of LHC Exotic Objects

### Displaced Leptons from $\tilde{\tau}_R$ s in Gauge Mediation

- HSCP Searches (CMS)

- Disappearing Tracks Searches (ATLAS & CMS)

- CMS Displaced  $e\mu$  Search

- Recast Limits

- Paths for Improvements

### A Gap: Same-Flavor Displaced Leptons

- Models: co-NLSPs, EGMSB, RPV, freezein dark matter

- A Same-Flavor Displaced Lepton Search for 13 TeV

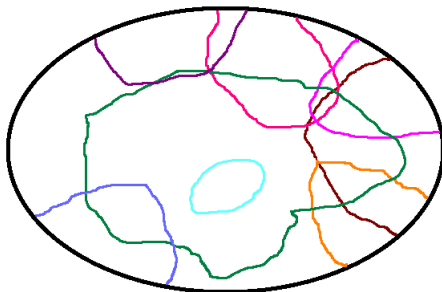
## Comments and Perspectives

# The LHC Program

The LHC has constrained many new particles in many models

- ▶ MSSM
- ▶  $t'/b'$
- ▶ UED
- ▶ GMSB
- ▶ RPV
- ▶ Stealth
- ▶ 2HDM
- ▶ ...

Signature Space



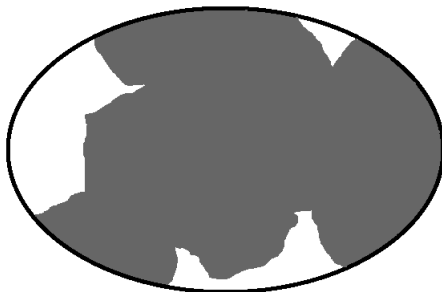
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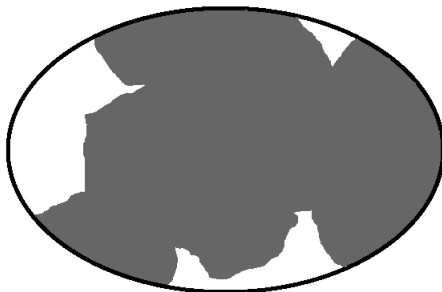
Need comprehensive program that covers **ALL** new physics scenarios

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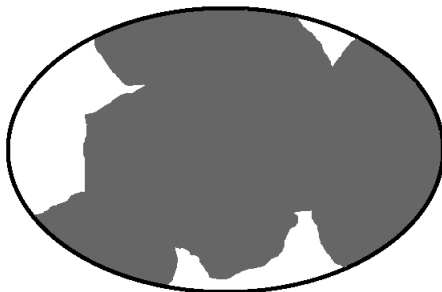
An **exotic object** could be our **FIRST** pathway to BSM physics!

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These searches cast a wide net – but there are **gaps**

Need comprehensive program that covers **ALL** new physics scenarios

An **exotic object** could be our **ONLY** pathway to BSM physics!!!

# What are Exotic Objects?

# What are Standard Objects?





# What are Exotic Objects?

<u>Object</u>	<u>Very Rough Identification Criteria</u>
1) Photon	Hard, isolated EM calo deposit, $E_{tracks} \ll E_{calo}$
2) Electron	Hard, isolated EM calo deposit, $E_{track} \sim E_{calo}$
3) Muon	Hard, isolated track through muon chamber
4) Jet	Other hard calo/track/particle clusters
a) Tau	Single or 3-prong hard, isolated track(s)
b) $b$ -jet	Secondary vertex, looks $b$ -ish
5) $\vec{E}_T$	$-\sum \vec{p}_T$

Loaded words: track, isolated, hard, cluster, vertex,  $b$ -ish ...

Exotic objects have properties that allow them to be distinguished from these standard objects

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Exotic objects have properties that allow them to be distinguished from these standard objects

Two basic classes: Direct & Indirect

# What are Exotic Objects?

Direct vs Indirect

Direct

Observe the object itself

Indirect

Observe atypical SM decay products

# What are Exotic Objects?

Direct vs Indirect

## Direct

Observe the object itself

Examples:

Disappearing tracks

Heavy, stable, charged particles

Magnetic monopoles

$R$ -hadrons

Quirks

...

## Indirect

Observe atypical SM decay products

# What are Exotic Objects?

Direct vs Indirect

## Direct

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LHC searches exist

## Indirect

Observe atypical SM decay products

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Direct vs Indirect

## Direct

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## Indirect

Observe atypical SM decay products

**Collimated particles fail isolation**

Non-isolated leptons/photons

Photon or lepton jets

# What are Exotic Objects?

Direct vs Indirect

## Direct

Observe the object itself

Examples:

Disappearing tracks

Heavy, stable, charged particles

Magnetic monopoles

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LHC searches exist

## Indirect

Observe atypical SM decay products

### Collimated particles fail isolation

Non-isolated leptons/photons

Photon or lepton jets

### Particles that decay in flight

Long lifetime from an approximate symmetry in the low energy theory

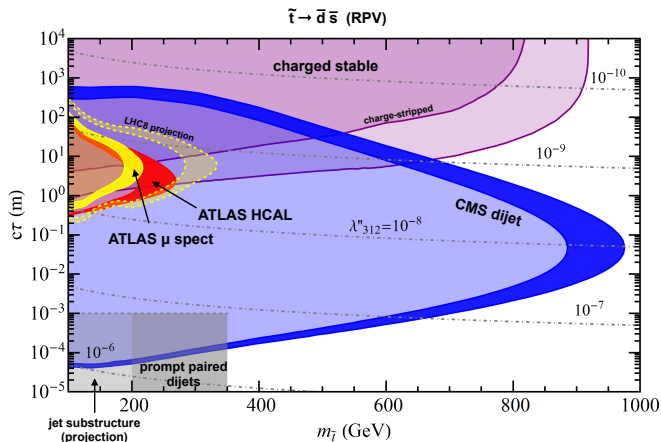
High dimension operators

High mass scale

Small couplings



# Long-lived / Displaced in SUSY



Liu, Tweedie – 2015

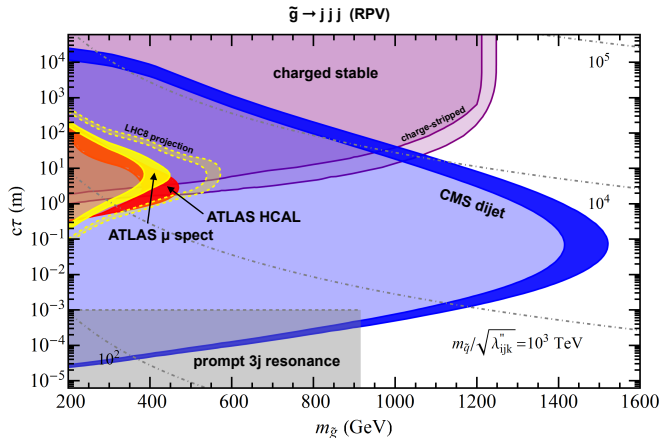
## Many displaced decays are well-covered: Most RPV

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Csaki, Kuflik, Lombardo, Slone, Volansky – 2015

Zwane – 2015

# Long-lived / Displaced in SUSY



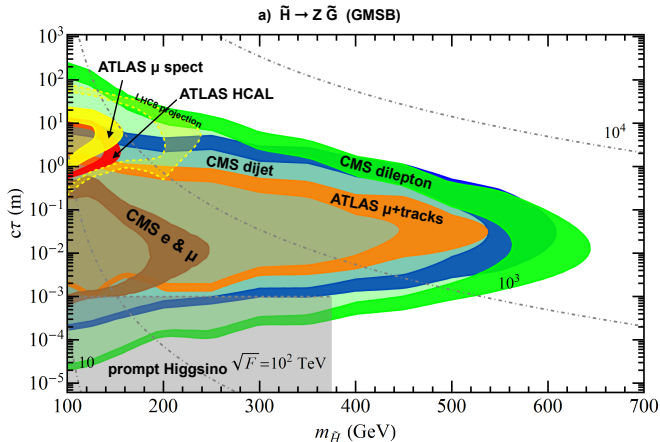
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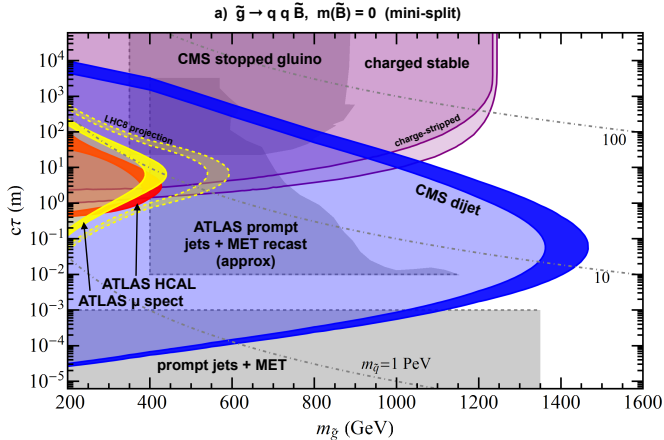


Liu, Tweedie – 2015

Many displaced decays are well-covered: Most GMSB

Liu, Tweedie – 2015

# Long-lived / Displaced in SUSY



Liu, Tweedie – 2015

Many displaced decays are well-covered: Mini-Split  
Coverage exceeds prompt signatures!

Liu, Tweedie – 2015

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$							
$b$	X						
$\gamma$	X	X					
$e$	X	X	X				
$\mu$	X	X	X	X			
$\tau$	X	X	X	X	X		
	Search			arXiv		Symbol	Comments

- ▶ Focused on pair produced, heavy decays inside the detector
- ▶ Only a selection of searches used, but fairly representative
- ▶ Cavalier about lifetime ranges and triggers; ignoring tops
- ▶ **Bold** is where searches are really optimized

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	jj	jj	jj	jj	jj	jj	jj
$b$	X	jj	jj	jj	jj	jj	jj
$\gamma$	X	X					
$e$	X	X	X				
$\mu$	X	X	X	X			
$\tau$	X	X	X	X	X		

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv
$b$	X	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv
$\gamma$	X	X					
$e$	X	X	X	dv	dv	dv	
$\mu$	X	X	X	X	dv	dv	
$\tau$	X	X	X	X	X	dv	

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV
ATLAS Displaced Vertex	1504.05162	dv	$m_X > 10$ GeV

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## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$E_T$
$j$	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv
$b$	X	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv	jj/dv
$\gamma$	X	X					
$e$	X	X	X	dv//	dv	dv//	
$\mu$	X	X	X	X	dv//	dv//	
$\tau$	X	X	X	X	X	dv//	

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV
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CMS Displaced Dilepton	1411.6977	//	$m_X > 15$ GeV



# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$E_T$
$j$	jj/dv	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv	jj/dv
$b$	X	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv	jj/dv
$\gamma$	X	X	dA	dA	dA	dA	dA
$e$	X	X	X	dv//	dv	dv//	
$\mu$	X	X	X	X	dv//	dv//	
$\tau$	X	X	X	X	X	dv//	

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV
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ATLAS Delayed Photon	1409.5542	dA	$m_X \gtrsim 100$ GeV

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	jj/dv	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
$b$	X	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
$\gamma$	X	X	dA	dA	dA	dA/em	dA
$e$	X	X	X	dv//	dv	dv//	
$\mu$	X	X	X	X	dv//	dv//	
$\tau$	X	X	X	X	X	dv//em	em

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV
ATLAS Displaced Vertex	1504.05162	dv	$m_X > 10$ GeV
CMS Displaced Dilepton	1411.6977	//	$m_X > 15$ GeV
ATLAS Delayed Photon	1409.5542	dA	$m_X \gtrsim 100$ GeV
CMS Displaced $e\mu$	1409.4789	em	Best at LFU

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Searches

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	jj/dv	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
$b$	X	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
$\gamma$	X	X	dA	dA	dA	dA/em	dA
$e$	X	X	X	dv//	dv	dv//	em
$\mu$	X	X	X	X	dv//	dv//	
$\tau$	X	X	X	X	X	dv//em	

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# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$							
$b$	X						
$\gamma$	X	X					
$e$	X	X	X				
$\mu$	X	X	X	X			
$\tau$	X	X	X	X	X		
	Models		Symbol				

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	<b>M</b>						<b>M</b>
$b$	X	<b>M</b>					<b>M</b>
$\gamma$	X	X					
$e$	X	X	X				
$\mu$	X	X	X	X			
$\tau$	X	X	X	X	X		
Models		Symbol					
Mini-Split		<b>M</b>					

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	MG						MG
$b$	X	MG					MG
$\gamma$	X	X					G
$e$	X	X	X	G			G
$\mu$	X	X	X	X	G		G
$\tau$	X	X	X	X	X	G	G

Models

Symbol

Mini-Split

M

GMSB

G

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	MGR	R		R	R	R	MGR
$b$	X	MGR		R	R	R	MGR
$\gamma$	X	X					G
$e$	X	X	X	GR	R	R	GR
$\mu$	X	X	X	X	GR	R	GR
$\tau$	X	X	X	X	X	GR	GR

Models

Symbol

Mini-Split

M

GMSB

G

RPV/dRPV

R

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	MGRS	R		R	R	R	MGR
$b$	X	MGRS		R	R	R	MGR
$\gamma$	X	X	S				G
$e$	X	X	X	GR	R	R	GR
$\mu$	X	X	X	X	GR	R	GR
$\tau$	X	X	X	X	X	GR	GR

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Stealth

S



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$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$\cancel{E}_T$
$j$	MGRS	R		R	R	R	MGR
$b$	X	MGRSH		R	R	R	MGR
$\gamma$	X	X	S				G
$e$	X	X	X	GR	R	R	GR
$\mu$	X	X	X	X	GRH	R	GR
$\tau$	X	X	X	X	X	GRH	GR

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Higgs Mixed

H

# Displaced Objects: $pp \rightarrow XX, X \rightarrow o_1 o_2(o_3)$

## Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$E_T$
$j$	MGRS $\gamma$	R		R	R	R	MGR
$b$	X	MGRSH $\gamma$		R	R	R	MGR
$\gamma$	X	X	S				G
$e$	X	X	X	GR $\gamma$	R	R	GR
$\mu$	X	X	X	X	GRH $\gamma$	R	GR
$\tau$	X	X	X	X	X	GRH $\gamma$	GR

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Higgs Mixed

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Dark Photon

$\gamma$

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## Models

$o_2 o_1$	$j$	$b$	$\gamma$	$e$	$\mu$	$\tau$	$E_T$
$j$	MGRS $\gamma$	R		R	R	R	MGRD
$b$	X	MGRSH $\gamma$		R	R	R	MGRD
$\gamma$	X	X	S				G
$e$	X	X	X	GR $\gamma$	R	R	GRD
$\mu$	X	X	X	X	GRH $\gamma$	R	GRD
$\tau$	X	X	X	X	X	GRH $\gamma$	GRD

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MD Freezein Dark Matter

D

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$e$	X	X	X	GR $\gamma$	R	R	GRD
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Models

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MD Freezein Dark Matter

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Well-motivated Theoretically

Weak Coverage Experimentally

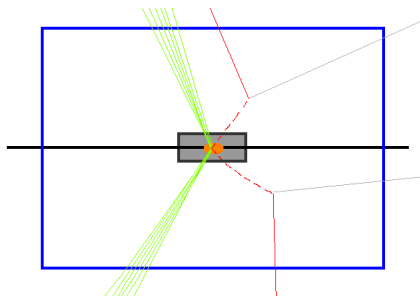
# Displaced Leptons in Prompt Searches

## Prompt lepton-based searches:

- ▶ Quality criteria drop displaced electrons
- ▶ Displaced muons veto events (cosmics)
- ▶ Vetoes range from  $50 \mu\text{m}$ – $1 \text{ mm}$

## Prompt jets+ $\cancel{E}_T$ searches:

- ▶ Veto events with leptons
- ▶ Definition not always transparent



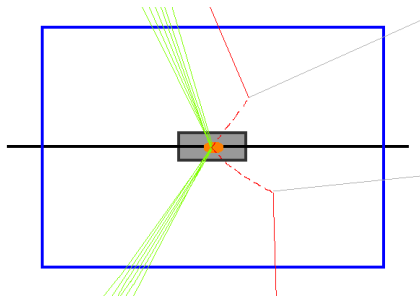
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## Prompt jets+ $\cancel{E}_T$ searches:

- ▶ Veto events with leptons
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Very dangerous region!

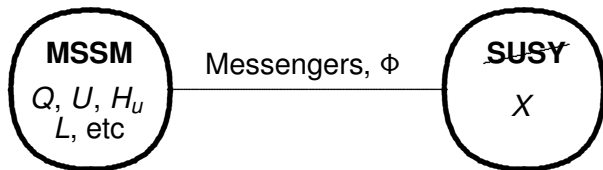
$$pp \rightarrow \tilde{\ell}^+ \tilde{\ell}^- + X \rightarrow \{\text{displaced muons}\} + X$$

lives in a prompt search blind spot!

Displaced electrons and taus  $\Rightarrow$  reduced efficiency

# Gauge Mediation and $\tilde{\tau}_R$ NLSPs

Lightning Review of Minimal GMSB

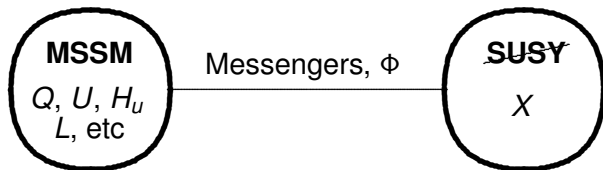


$$W \sim X\phi\tilde{\phi} + \{\text{MSSM yukawas}\}$$

$$\langle X \rangle = M + \theta^2 F, \quad \Lambda \equiv F/M, \quad \tilde{\Lambda} \equiv \frac{\Lambda}{16\pi^2}$$

# Gauge Mediation and $\tilde{\tau}_R$ NLSPs

## Lightning Review of Minimal GMSB



$$W \sim X\phi\tilde{\phi} + \{\text{MSSM yukawas}\}$$

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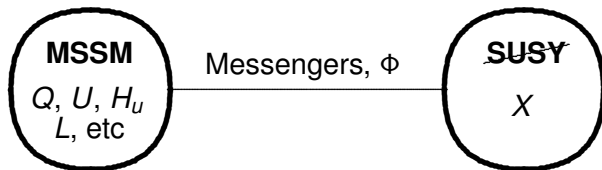
$$M_r \sim N_{\text{eff}} g_r^2 \tilde{\Lambda} \quad A\text{-terms} = 0$$

$$m_{\text{soft}}^2 \sim 2N_{\text{eff}} C_r g_r^4 \tilde{\Lambda}^2 \quad (C_r \text{ quadratic Casimirs } \mathcal{O}(1))$$



# Gauge Mediation and $\tilde{t}_R$ NLSPs

## Lightning Review of Minimal GMSB



$$W \sim X\phi\tilde{\phi} + \{\text{MSSM yukawas}\}$$

$$\langle X \rangle = M + \theta^2 F, \quad \Lambda \equiv F/M, \quad \tilde{\Lambda} \equiv \frac{\Lambda}{16\pi^2}$$

$$M_r \sim N_{\text{eff}} g_r^2 \tilde{\Lambda} \quad A\text{-terms} = 0$$

$$m_{\text{soft}}^2 \sim 2N_{\text{eff}} C_r g_r^4 \tilde{\Lambda}^2 \quad (C_r \text{ quadratic Casimirs } \mathcal{O}(1))$$

Potential NLSP Masses:  $\begin{cases} m_{\tilde{B}} = N_{\text{eff}} g_1^2 \tilde{\Lambda} \\ m_{\tilde{\ell}_R} = \sqrt{\frac{6N_{\text{eff}}}{5}} g_1^2 \tilde{\Lambda} \end{cases} \quad N_{\text{eff}} \geq 2 \Rightarrow \tilde{t}_R \text{ NLSP}$   
(or large running)

GMSB is a very **well-motivated** source of displaced particles

$$c\tau \approx 100 \mu\text{m} \left( \frac{100 \text{ GeV}}{m_{\tilde{\tau}}} \right)^5 \left( \frac{\sqrt{F}}{100 \text{ TeV}} \right)^4$$

What is  $\sqrt{F}$ ?

# Gauge Mediation and $\tilde{\tau}_R$ NLSPs

Lifetimes

GMSB is a very **well-motivated** source of displaced particles

$$c\tau \approx 100 \mu\text{m} \left( \frac{100 \text{ GeV}}{m_{\tilde{\tau}}} \right)^5 \left( \frac{\sqrt{F}}{100 \text{ TeV}} \right)^4$$

What is  $\sqrt{F}$ ?

$F < M^2$ ; otherwise arbitrary

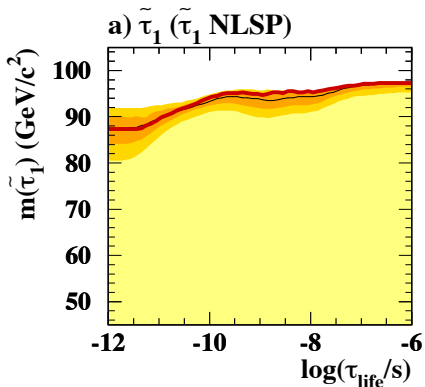
$$c\tau \sim 10 \mu\text{m} \left( \frac{100 \text{ GeV}}{m_{\tilde{\tau}}} \right) \left( \frac{M}{\sqrt{F}} \right)^4 \frac{1}{N_{\text{eff}}^2} \quad (\text{minimal GM only})$$

LHC relevant range:  $100 \mu\text{m} \lesssim c\tau \lesssim 1 \text{ m}$

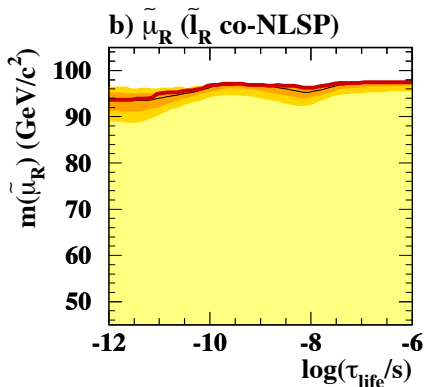
Measuring  $m_{\tilde{\tau}_R}$  &  $c\tau_{\tilde{\tau}_R}$  probes **SUSY breaking!**

# Gauge Mediation and $\tilde{\tau}_R$ NLSPs

LEP Limits on Slepton NLSPs



$$m_{\tilde{\tau}} > 87 \text{ GeV}$$



$$m_{\tilde{\mu}} > 94 \text{ GeV}$$

OPAL placed the best limits on sleptons of all lifetimes

# Relevant LHC search: HSCP

At long lifetime:  $c\tau_{\tilde{\tau}} \gtrsim 1\text{m} \Rightarrow$  heavy, detector-stable, charged particle

CMS HSCP search 1305.0491 (expect similar from ATLAS 1411.6795)

## Cuts

Central:  $|\eta| < 2.4$

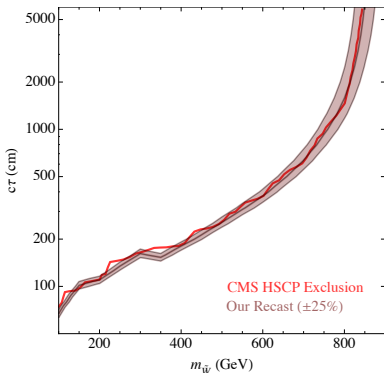
Hard:  $p_T > 70\text{ GeV}$

Isolated:  $\sum p_{T,\Delta R < 0.3}^{trks} < 50\text{ GeV}$   
 $p_{\Delta R < 0.3}^{calo, track} < 0.3$

Slow:  $1/\beta > 1.225$

Large  $dE/dx$ : see 1305.0491

Efficiency maps provided: 1502.02522



# Relevant LHC search: Disappearing Tracks

CMS 1411.6006

At slightly shorter lifetimes:  $30 \text{ cm} \lesssim c\tau_{\tilde{\tau}} \lesssim 3 \text{ m} \Rightarrow$  disappearing tracks

## Cuts

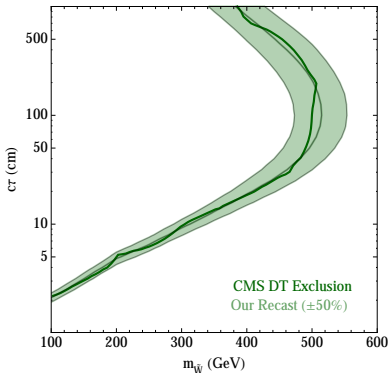
A hard jet:  $p_T > 110 \text{ GeV}$ ,  $|\eta| < 2.4$

Missing energy:  $\cancel{E}_T > 100 \text{ GeV}$

Hard track stub:  
 $p_T > 50 \text{ GeV}$   
Few outer tracker hits

Particular eta:  $1.85 < |\eta| < 2.1$   
 $0.35 < |\eta| < 1.42$   
or  $|\eta| < 0.15$

Isolated:  $E_{\Delta R < 0.5}^{calo} < 10 \text{ GeV}$   
 $I_{\Delta R < 0.3}^{trks, track} < 0.05$



Simple efficiency map provided 1411.6006

# Relevant LHC search: Disappearing Tracks

ATLAS 1310.3675

At slightly shorter lifetimes:  $20 \text{ cm} \lesssim c\tau_{\tilde{\tau}} \lesssim 2 \text{ m} \Rightarrow$  disappearing tracks

## Cuts

A hard jet:  $p_T > 90 \text{ GeV}$ ,  $|\eta| < 2.5$

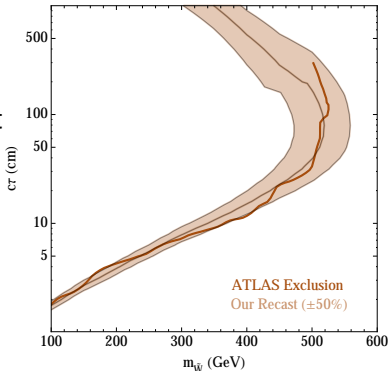
Missing energy:  $\cancel{E}_T > 90 \text{ GeV}$

Hard track stub:  
 $p_T > 75 \text{ GeV}$   
highest  $p_T$  track in event  
< 5 hits in TRT

Eta:  $0.1 < |\eta| < 1.9$

Isolated:  $I_{\Delta R < 0.4}^{trks, track} < 0.04$   
 $\Delta R_{jt} > 0.4 \forall p_{T,j} > 45 \text{ GeV}$

Veto leptons:  $e_s, \mu_s$ , &  
 $\mu$  calo tracks  $> 10 \text{ GeV}$



No efficiency map provided

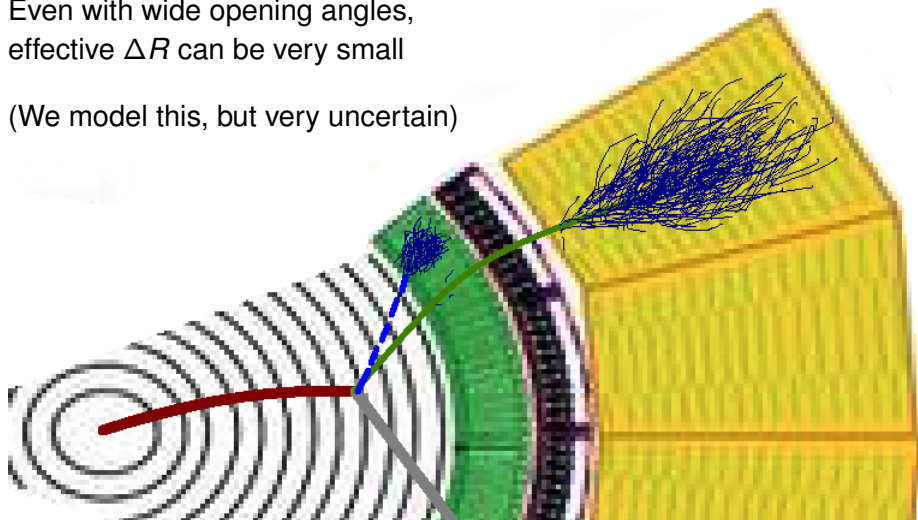
# Relevant LHC search: Disappearing Tracks

## Recast Caveats

Isolation requirements are hard on long-lived staus

Even with wide opening angles,  
effective  $\Delta R$  can be very small

(We model this, but very uncertain)





# Relevant LHC search: CMS Displaced $e\mu$ (1409.4789)

## Cuts

### Cut Summary of CMS $e\mu$

#### Preselection

1 OS  $e^\pm\mu^\mp$  pair

→  $d_\ell > 100 \mu\text{m}$

$p_{T,\ell} > 25 \text{ GeV}$ ,  $|\eta_\ell| < 2.5$

Reject  $1.44 < |\eta_e| < 1.56$

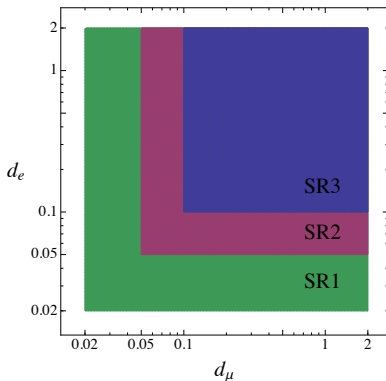
$I_{\Delta R < 0.3}^{calo,e} < 0.10$ ,  $I_{\Delta R < 0.4}^{calo,\mu} < 0.12$

$\Delta R_{\ell j} > 0.5 \forall \text{ jets with } p_T > 10 \text{ GeV}$

$\Delta R_{e\mu} > 0.5$

$v_{T,\tilde{\ell}} < 4 \text{ cm}$ ,  $v_{Z,\tilde{\ell}} < 30 \text{ cm}$

Veto additional leptons



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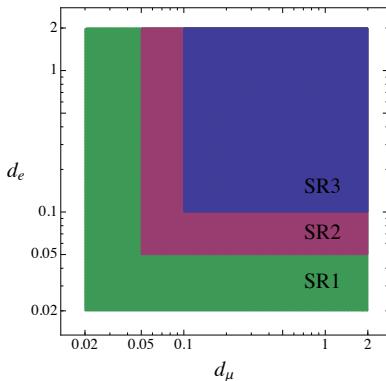
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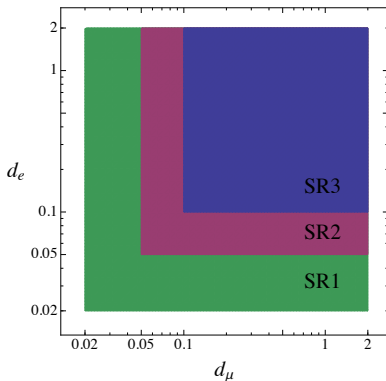
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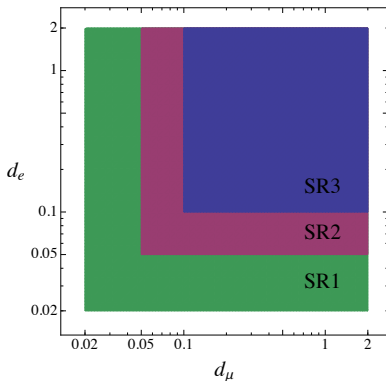
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$\Delta R_{ej} > 0.5 \forall$  jets with  $p_T > 10 \text{ GeV}$

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→ Veto additional leptons



# Relevant LHC search: CMS Displaced $e\mu$ (1409.4789)

Recast

## Cut Summary of CMS $e\mu$

### Preselection

1 OS  $e^\pm\mu^\mp$  pair

$d_\ell > 100 \mu\text{m}$

$p_{T,\ell} > 25 \text{ GeV}$ ,  $|\eta_\ell| < 2.5$

Reject  $1.44 < |\eta_e| < 1.56$

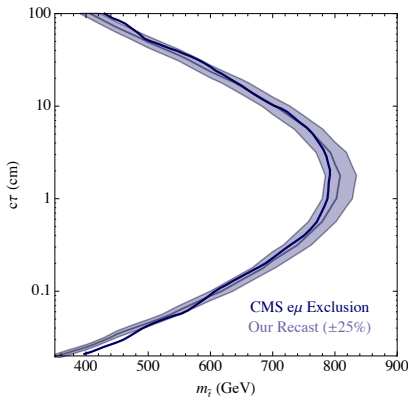
$I_{\Delta R < 0.3}^{\text{calo},e} < 0.10$ ,  $I_{\Delta R < 0.4}^{\text{calo},\mu} < 0.12$

$\Delta R_{\ell j} > 0.5 \forall \text{ jets with } p_T > 10 \text{ GeV}$

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Veto additional leptons

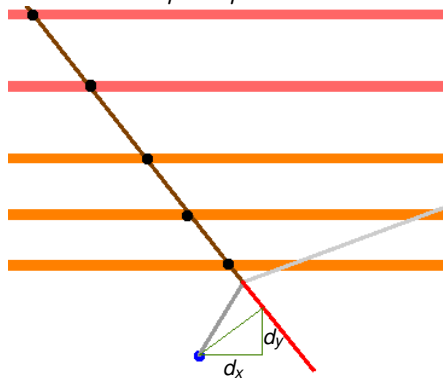


Extensive recasting details provided!

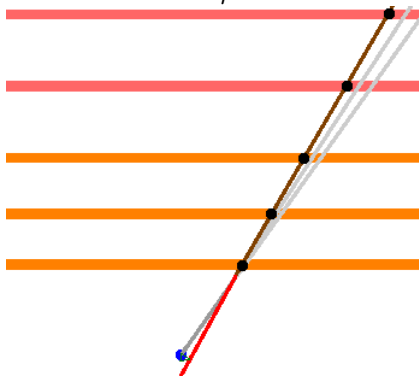
# Relevant LHC search: CMS Displaced $e\mu$ (1409.4789)

Impact Parameter

$$\tilde{\mu} \rightarrow \mu \tilde{G}$$



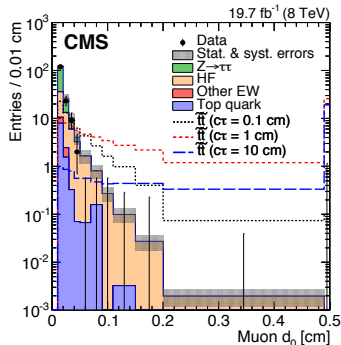
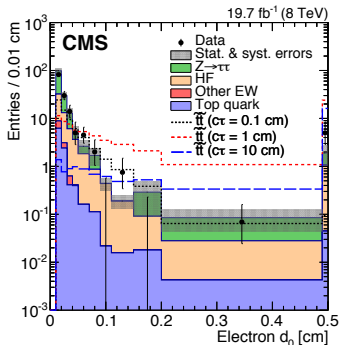
$$\tau \rightarrow \mu \nu \bar{\nu}$$



Impact Parameter is *not* the location of parent  $b$  and  $\tau$  decay products are more collimated

# Relevant LHC search: CMS Displaced $e\mu$ (1409.4789)

## Backgrounds & Data



Event source	SR1	SR2	SR3
Other EW	$0.65 \pm 0.13 \pm 0.09$	$(0.89 \pm 0.53 \pm 0.12) \times 10^{-2}$	$< (89 \pm 53 \pm 12) \times 10^{-4}$
Top quark	$0.77 \pm 0.04 \pm 0.08$	$(1.25 \pm 0.26 \pm 0.12) \times 10^{-2}$	$(2.4 \pm 1.3 \pm 0.2) \times 10^{-4}$
Z $\rightarrow$ $\tau\tau$	$3.93 \pm 0.42 \pm 0.39$	$(0.73 \pm 0.73 \pm 0.07) \times 10^{-2}$	$< (73 \pm 73 \pm 7) \times 10^{-4}$
HF	$12.7 \pm 0.2 \pm 3.8$	$(98 \pm 6 \pm 30) \times 10^{-2}$	$(340 \pm 110 \pm 100) \times 10^{-4}$
Total expected background	$18.0 \pm 0.5 \pm 3.8$	$1.01 \pm 0.06 \pm 0.30$	$0.051 \pm 0.015 \pm 0.010$
Observed	19	0	0

Only HSCP limits on direct  $\tilde{\tau}_R$  production!



# Recast Limits on $\tilde{\tau}_R$

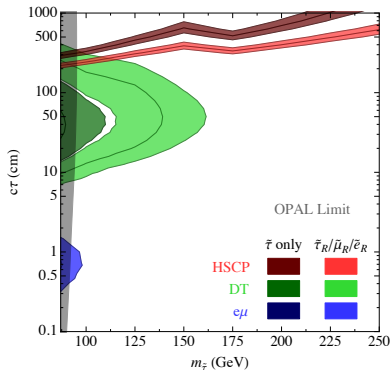
Only HSCP limits on direct  $\tilde{\tau}_R$  production!

But . . . a  $\tilde{\tau}_R$  is not expected in isolation

Near degenerate slepton limits

$$m_{\tilde{e}_R} = m_{\tilde{\mu}_R} = m_{\tilde{\tau}_R} + 10 \text{ GeV}$$

$$\tilde{\ell}_R \rightarrow \tilde{\tau}_R + \{\text{soft}\}$$



(Better limit from both disappearing track searches shown)

# Recast Limits on $\tilde{\tau}_R$

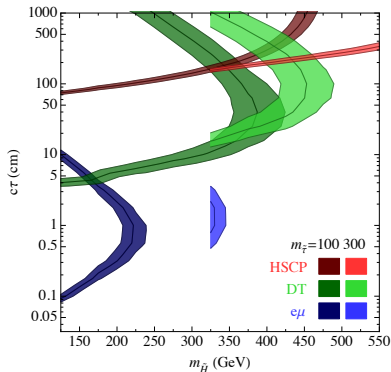
Only HSCP limits on direct  $\tilde{\tau}_R$  production!

But . . . a  $\tilde{\tau}_R$  is not expected in isolation

## Higgsino production limits

$$m_{\tilde{\chi}_1^0} = m_{\tilde{\chi}_2^0} = m_{\tilde{\chi}_1^\pm}$$

$$\tilde{\chi}_{1,2}^0 \rightarrow \tilde{\tau}_R^\pm \tau^\mp, \quad \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}_R^\pm \nu$$



(Better limit from both disappearing track searches shown)

Limits are very sensitive to  $m_{\tilde{\tau}_R}$

# Recast Limits on $\tilde{\tau}_R$

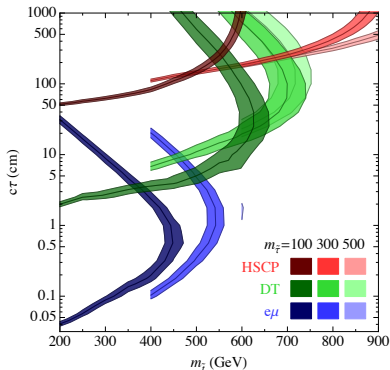
Only HSCP limits on direct  $\tilde{\tau}_R$  production!

But . . . a  $\tilde{\tau}_R$  is not expected in isolation

Stop production limits

$$m_{\tilde{H}} = m_{\tilde{t}} - 50 \text{ GeV}$$

$$\tilde{t} \rightarrow b\tilde{H}^+ \rightarrow b\nu\tilde{\tau}_R^+$$



(Better limit from both disappearing track searches shown)

Limits are very sensitive to  $m_{\tilde{\tau}_R}$

# Recast Limits on $\tilde{\tau}_R$

Only HSCP limits on direct  $\tilde{\tau}_R$  production!

But . . . a  $\tilde{\tau}_R$  is not expected in isolation

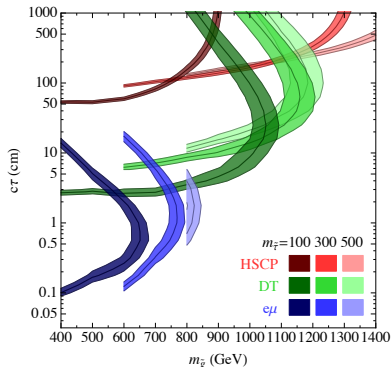
## Glino production limits

$$m_{\tilde{t}} = m_{\tilde{g}} - 200 \text{ GeV}$$

$$m_{\tilde{H}} = m_{\tilde{t}} - 50 \text{ GeV}$$

$$\tilde{g} \rightarrow \tilde{t}\bar{t} \rightarrow \bar{t}b\tilde{H}^+ \rightarrow \bar{t}b\nu\tilde{\tau}_R^+$$

$$\tilde{g} \rightarrow \tilde{t}^*t \rightarrow t\bar{b}\tilde{H}^- \rightarrow t\bar{b}\bar{\nu}\tilde{\tau}_R^-$$



(Better limit from both disappearing track searches shown)

Limits are very sensitive to  $m_{\tilde{\tau}_R}$

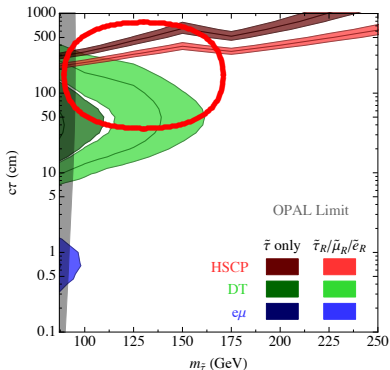
# Potential Improvements?

HSCP

CMS heavy, stable, charged particle search does very well!

Using “tracker-only” signal region could improve sensitivity

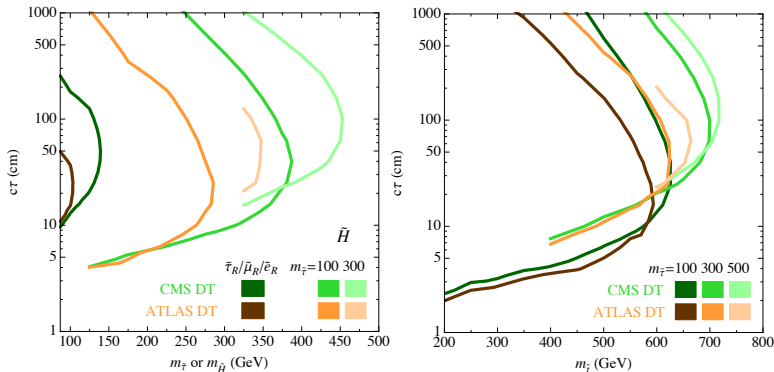
Currently no efficiency maps provided



# Potential Improvements?

## Disappearing Tracks

ATLAS hurt by lepton vetoes and hardest track requirement



Both disappearing tracks searches are hurt by isolation requirements

Could similar pre-selection capitalize on kinked track?

# Potential Improvements?

## CMS Displaced Lepton Search

There are several lessons from GMSB  $\tilde{\tau}_R$ s to improve sensitivity

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow e^\pm \mu^\mp + X) = 6\%$$

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow e^+ e^- + X) = 3\%$$

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow \mu^+ \mu^- + X) = 3\%$$

1) Add same-flavor lepton channels

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1) Add same-flavor lepton channels

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow e^\pm \tau_h^\mp + X) = 23\%$$

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow \mu^\pm \tau_h^\mp + X) = 23\%$$

$$BR(\tilde{\tau}^+ \tilde{\tau}^- \rightarrow \tau_h^\pm \tau_h^\mp + X) = 42\%$$

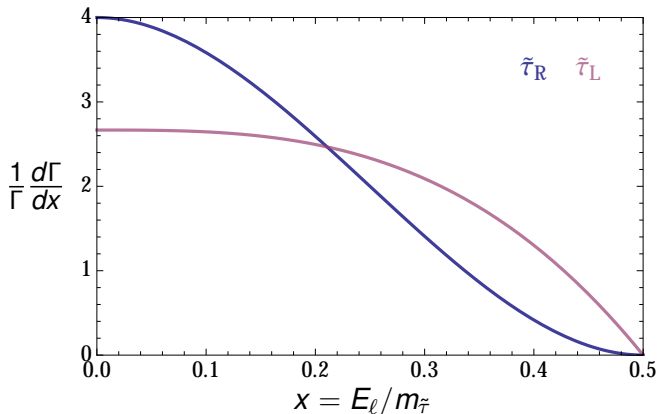
2) Include hadronic  $\tau_h$ s

Experimental feasibility of displaced  $\tau_h$ s?



# Potential Improvements?

## CMS Displaced Lepton Search



Right-handed polarized  $\tau$ s from  $\tilde{\tau}_R$  decays give softer leptons

3) Lower  $p_T$  thresholds can capture a lot more signal

Additional triggers –  $\cancel{E}_T + ll$ ,  $\cancel{E}_T, lll$ , etc

# Potential Improvements?

CMS Displaced Lepton Search

Search vetoes additional leptons.

Why?

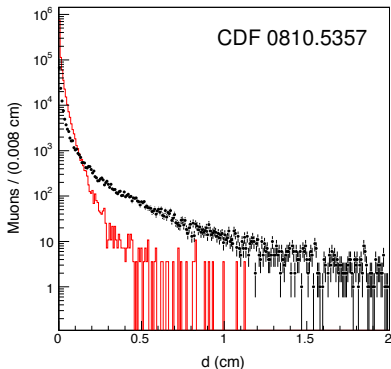
Displaced multilepton background  
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(CDF ghost muons???)

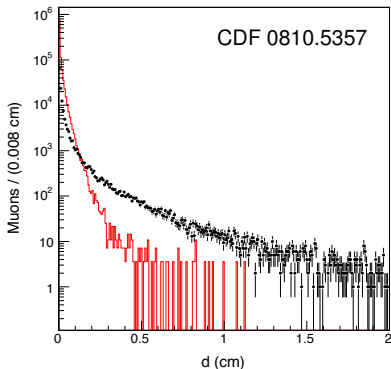


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CMS Displaced Lepton Search

Search vetoes additional leptons.  
Why?

Displaced multilepton background  
should be very small  
(CDF ghost muons???)



Glauino & Higgsino model have additional leptons often  $\sim (45\%, 30\%)$

If pair-produced object is not charged under lepton number,  
additional leptons are generic

4) Don't veto additional leptons

# Potential Improvements?

CMS Displaced Lepton Search

Gluino & Higgsino models have Majorana particles in chain

⇒ same-sign displaced leptons

5) Include same-sign displaced lepton signal regions

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CMS Displaced Lepton Search

Gluino & Higgsino models have Majorana particles in chain

⇒ same-sign displaced leptons

5) Include same-sign displaced lepton signal regions

5') Same-sign possibility fairly generic, be wary of CR contamination

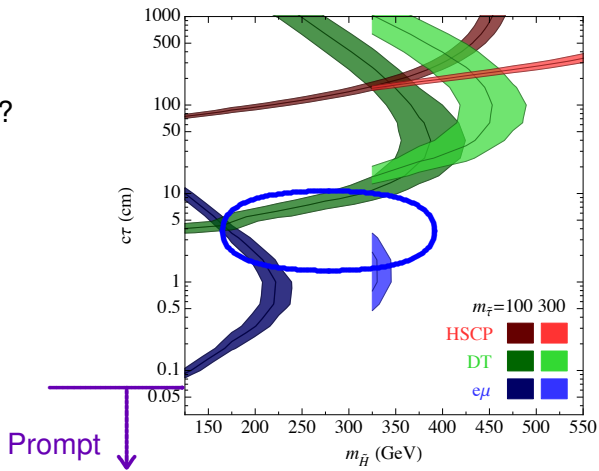
$SS\ell$  can appear in the  $\tilde{t} \rightarrow \ell^+ b$  benchmark of CMS 1409.4789

Mesino oscillation allows up to 3/8 of events as  $SS\ell$  Sarid, Thomas – 9909349

# Potential Improvements?

## CMS Displaced Lepton Search

Extend reach in  $c\tau$ ?

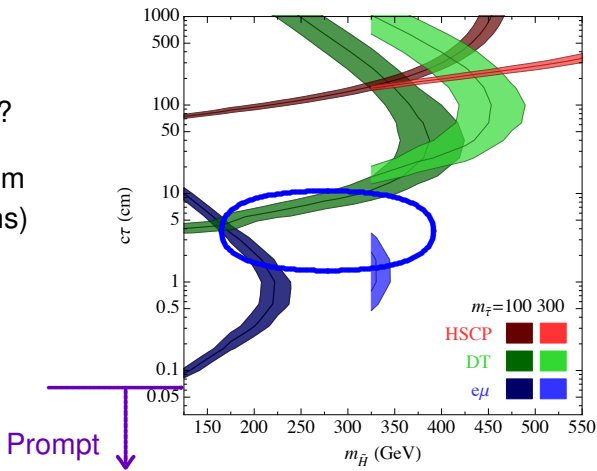


# Potential Improvements?

## CMS Displaced Lepton Search

Extend reach in  $c\tau$ ?

- 6) Allow  $d_0$  above 2 cm  
(Even just for muons)





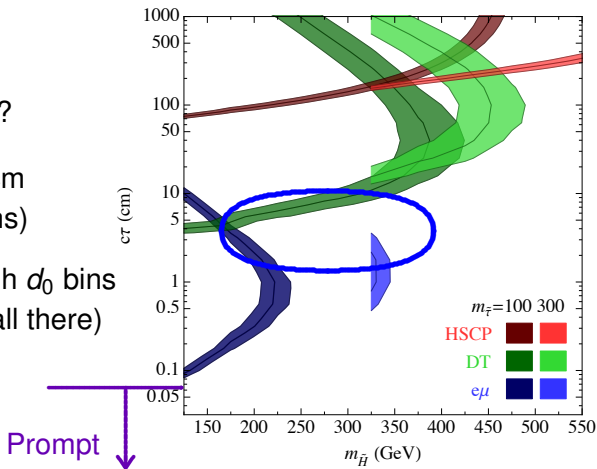
# Potential Improvements?

## CMS Displaced Lepton Search

Extend reach in  $c\tau$ ?

6) Allow  $d_0$  above 2 cm  
(Even just for muons)

7) Relax isolation in high  $d_0$  bins  
(Backgrounds are small there)



# Same-flavor Displaced Lepton Models

Slepton Co-NLSP in GGM

If  $m_{\tilde{\ell}_R} - m_{\tilde{\tau}_R} \ll 10 \text{ GeV}$       or       $m_{\tilde{B}} \gg m_{\tilde{\tau}_R}$ ,

then  $\Gamma(\tilde{\ell}_R \rightarrow \ell \tau \tilde{\tau}_R) \ll \Gamma(\tilde{\ell}_R \rightarrow \ell \tilde{G}) \Rightarrow$  Slepton Co-NLSP

Events have displaced  $e^+ e^-$ ,  $\mu^+ \mu^-$ , or  $\tau^+ \tau^-$

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Small splitting can happen for low  $\tan \beta$

In GGM,  $M_1$  and  $m_{\tilde{E}_R}^2$  are independent

$\mu^+ \mu^-$  and  $e^+ e^-$  searches would be more sensitive to this model

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Can one get 100%  $e^+ e^-$  or  $\mu^+ \mu^-$  without  $\tau^+ \tau^-$ ?

# Same-flavor Displaced Lepton Models

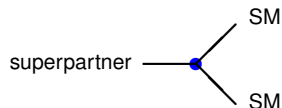
$\tilde{\tau}_1$  with *LLE* RPV

RPV can do almost anything

# Same-flavor Displaced Lepton Models

$\tilde{\tau}_1$  with *LLE* RPV

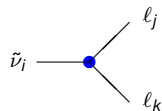
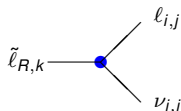
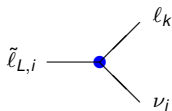
RPV can do almost anything



$$W = \frac{1}{2} \lambda_{ijk} L_i L_j E_k^c + \lambda'_{ijk} L_i Q_j D_k^c + \frac{1}{2} \lambda''_{ijk} U_i^c D_j^c D_k^c + \mu_i L_i H_U$$

$i, j, k$  = generation indices

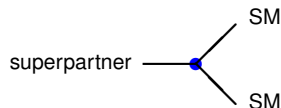
*LLE*



# Same-flavor Displaced Lepton Models

$\tilde{\tau}_1$  with *LLE* RPV

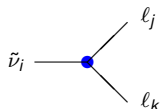
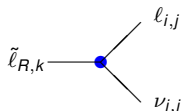
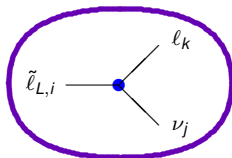
RPV can do almost anything



$$W = \frac{1}{2} \lambda_{ijk} L_i L_j E_k^c + \lambda'_{ijk} L_i Q_j D_k^c + \frac{1}{2} \lambda''_{ijk} U_i^c D_j^c D_k^c + \mu_i L_i H_U$$

$i, j, k =$  generation indices

*LLE*



$$c_T \approx 1 \text{ cm} \left( \frac{10^{-7}}{\lambda_{232}} \right)^2 \left( \frac{100 \text{ GeV}}{m_{\tilde{\tau}}} \right)^2 \text{ sec}^2 \theta_{\tilde{\tau}} \quad (\theta_{\tilde{\tau}} = 0 \Rightarrow \tilde{\tau}_1 = \tilde{\tau}_L)$$

$$\lambda_{232} \cos \theta_{\tilde{\tau}} \gg \text{other RPV} \Rightarrow \text{BR}(\tilde{\tau}_1 \rightarrow \mu \nu) \approx 100\%$$

# Same-flavor Displaced Lepton Models

EGMSB

Extended GMSB (EGMSB) models can generate large  $A_t$   
for maximal  $\tilde{t}$  mixing and decreased tuning of the MSSM Higgs

Craig, Knapen, Shih, Zhao – 1206.4086; Evans, Shih – 1303.0228; others

EGMSB can *also* give a 1st- or 2nd-gen slepton NLSP Shadmi, Szabo – 1103.0292



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Add EGMSB MSSM-Messenger coupling:  $W \supset \kappa_i E_i^c \Phi_U \Phi_{\bar{D}}$

$$\vec{\kappa} = (0, \kappa_2, 0) \text{ gives } \Delta m_{\tilde{\mu}} \sim -25\kappa_2^2 m_{\tilde{\ell}}$$

$$\kappa_2 \sim 0.1 \Rightarrow \tilde{\mu}_R \text{ NLSP}$$

$\frac{\kappa_1}{\kappa_2} \ll 0.1$  &  $\frac{\kappa_3}{\kappa_2} \lesssim 0.3$  no CMS  $e\mu$  sensitivity and safe from flavor

# Same-flavor Displaced Lepton Models

Freezein of Dark Matter During a Matter-Dominated Era

Minimal models of lepton-flavored DM

Model	Mediator	DM	Lagrangian
1	Scalar	Fermion	$y_{ij}^{LDM} \ell_i^c \zeta^- \chi_i + m_\zeta^2 \zeta^+ \zeta^- + m_{\chi,ij} \chi_i \bar{\chi}_j$
2	Fermion	Scalar	$y_{ij}^{LDM} \ell_i^c \psi \mathbf{S}_j + m_\psi \psi \bar{\psi} + m_{\mathbf{S},ij}^2 \mathbf{S}_i^\dagger \mathbf{S}_j$

$y^{LDM}$  for  $c_T \sim \mathcal{O}(1 \text{ mm} - 1 \text{ m}) \Rightarrow \Omega_{DM}$  too high!

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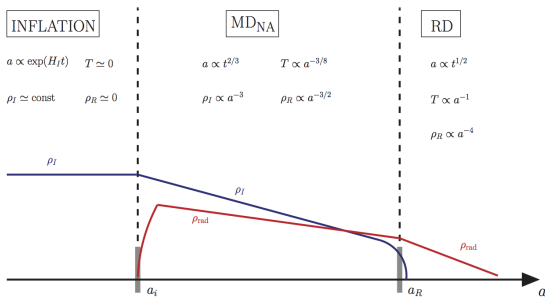
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Inflaton decay  
 $\Downarrow$   
 early matter-dominated era

$T_{RH} < m_\psi / \zeta \Rightarrow$  less  $\Omega_{DM}$   
 due to entropy injection



Co, D'Eramo, Hall, Pappadopulo - 1506.07532

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For Model 2:

$$\frac{\Omega_{DM}}{\Omega_{DM,obs}} \approx \left( \frac{20}{m_\psi / T_{RH}} \right)^7 \left( \frac{500 \text{ GeV}}{m_\psi} \right)^2 \left( \frac{m_S}{1 \text{ MeV}} \right) \left( \frac{1 \text{ cm}}{c\tau_\psi} \right)$$

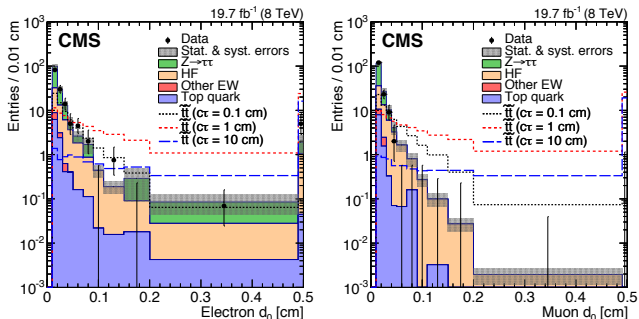
For  $100 \text{ GeV} < m_\psi \lesssim 1 \text{ TeV}$  and  $100 \mu\text{m} \lesssim c\tau \lesssim 1 \text{ m}$

$m_S \ll m_\psi$  and  $T_{RH}$  can be chosen to give  $\Omega_{DM} = \Omega_{DM,obs}$ !

# Same-Flavor Displaced Lepton Search

## Backgrounds

How can we estimate 13 TeV backgrounds?



1. Use 8 TeV backgrounds to estimate 8 TeV SF backgrounds
2. Rescale backgrounds by  $\frac{\sigma(13)}{\sigma(8)}$  (supported by HF MC)
3. Assume displaced  $Z \rightarrow e^+ e^- / \mu^+ \mu^-$  is small or can be controlled
4. Artificially low trigger, but have a background we can estimate

# Same-Flavor Displaced Lepton Search

## Backgrounds

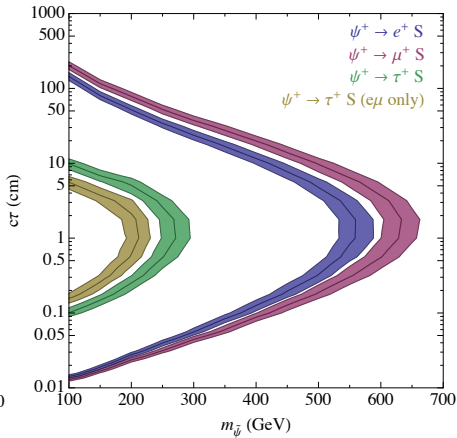
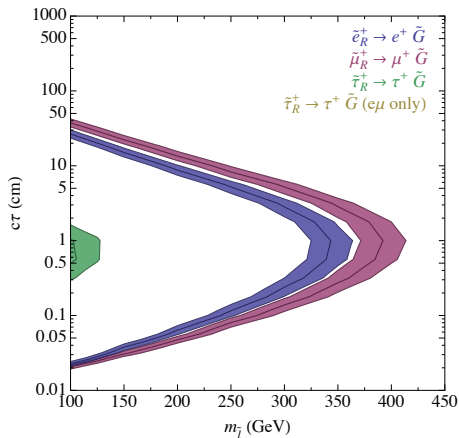
How can we estimate 13 TeV backgrounds?

Sample	SR1	SR2	SR3
$e^\pm \mu^\mp$ 8 TeV (CMS actual)	$18.0 \pm 3.8$	$1.01 \pm 0.31$	$0.051 \pm 0.018$
$e^\pm \mu^\mp$ 8 TeV (estimate)	$19.8 \pm 4.1$	$0.92 \pm 0.28$	$0.055 \pm 0.024$
$e^\pm \mu^\mp$ 13 TeV ( $20 \text{ fb}^{-1}$ )	$34.1 \pm 6.5$	$1.49 \pm 0.44$	$0.086 \pm 0.038$
$e^+ e^-$ 13 TeV ( $20 \text{ fb}^{-1}$ )	$25.2 \pm 3.6$	$1.43 \pm 0.33$	$0.31 \pm 0.06$
$\mu^+ \mu^-$ 13 TeV ( $20 \text{ fb}^{-1}$ )	$13.0 \pm 3.1$	$0.50 \pm 0.15$	$0.012 \pm 0.006$

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# Same-Flavor Displaced Lepton Search

13 TeV Same-Flavor Search



# Same-Flavor Displaced Lepton Search

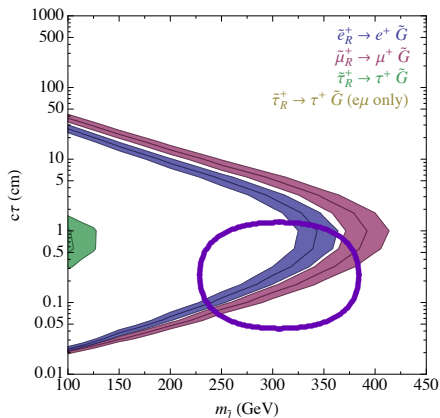
Improving a 13 TeV Same-Flavor Search

$\tilde{e}^+ \rightarrow e^+ \tilde{G} \Rightarrow$  hard leptons

$\tilde{\mu}^+ \rightarrow \mu^+ \tilde{G} \Rightarrow$  hard leptons

$\tilde{\tau}^+ \rightarrow \tau^+ \tilde{G} \Rightarrow$  soft leptons

8) SF search can be improved with  
higher  $p_{T,\ell}$ , lower background bins

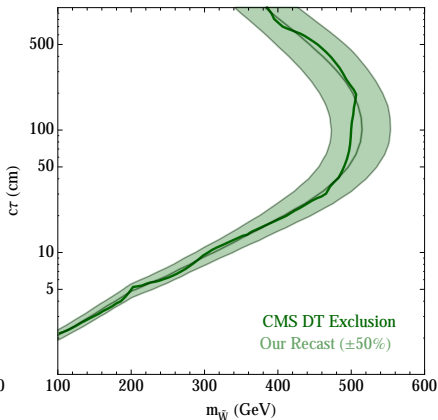
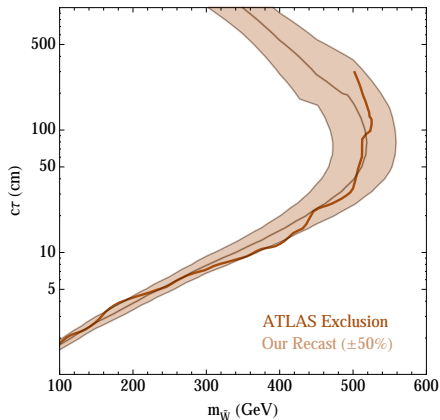




# Comments and Perspectives

## Recasting

Efficiency maps and recasting instructions are an *essential* facet to all searches for exotic objects

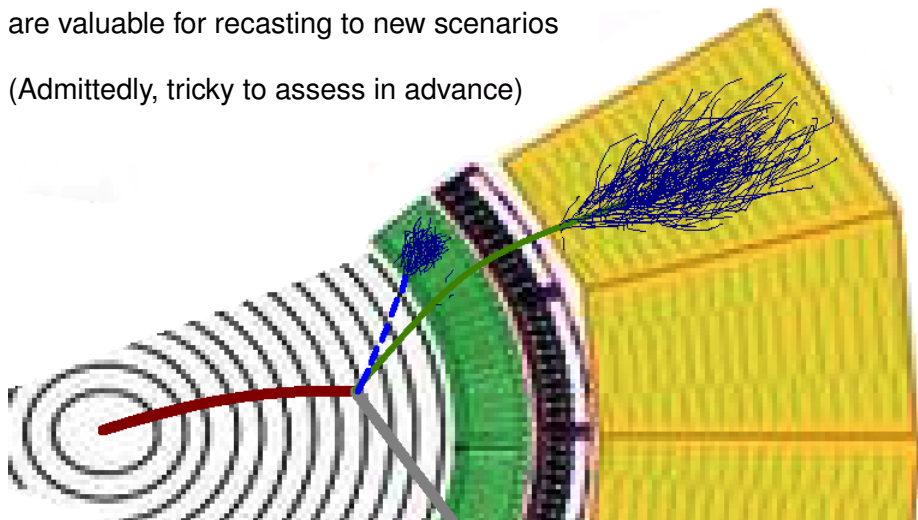


# Comments and Perspectives

## Recasting

Clear information about applying the search **beyond** the benchmark  
are valuable for recasting to new scenarios

(Admittedly, tricky to assess in advance)



# Comments and Perspectives

Known unknowns. . .

What more is wanted from **theory**?

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What regions of parameter space ( $m_Q$  vs  $\Lambda_{IC}$ ) are constrained?

What new search strategies could fill the gaps?

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- ▶ Quirks  
What regions of parameter space ( $m_Q$  vs  $\Lambda_{IC}$ ) are constrained?  
What new search strategies could fill the gaps?
- ▶ Hidden valleys (high mass and Higgs portal)  
What classes of models are constrained by existing searches?  
Can model-specific details be distilled to a simplified framework?  
Minimal set of searches to cover all observable possibilities?

# Comments and Perspectives

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- ▶ Unknown unknowns    {insert your paper here}

# Conclusions

- ▶ Many exotic objects are covered, but there are gaps
- ▶ Displaced  $\tilde{\tau}$ s from GMSB: well-motivated and weakly constrained
- ▶ Sensitivity to  $\tilde{\tau}_R$  can be improved in the CMS  $e^\pm\mu^\mp$  search
  - ▶ Add SF $\ell$  bins
  - ▶ Add  $\tau_h$  bins
  - ▶ Lowered  $p_T$  thresholds
  - ▶ Extend  $d_0 > 2$  cm
  - ▶ Add SS $\ell$  bins (CR contamination)
  - ▶ Allow extra  $\ell$ s
  - ▶ Relax isolation in high  $d_0$  bins
  - ▶ Add high  $p_{T,\ell}$  bins
- ▶ Disappearing track searches should consider this benchmark
  - ▶ Kinked tracks?
- ▶ Several models with displaced  $ee/\mu\mu$  uncovered at LHC
- ▶ A lot of exciting work needs to be done on exotic objects!