

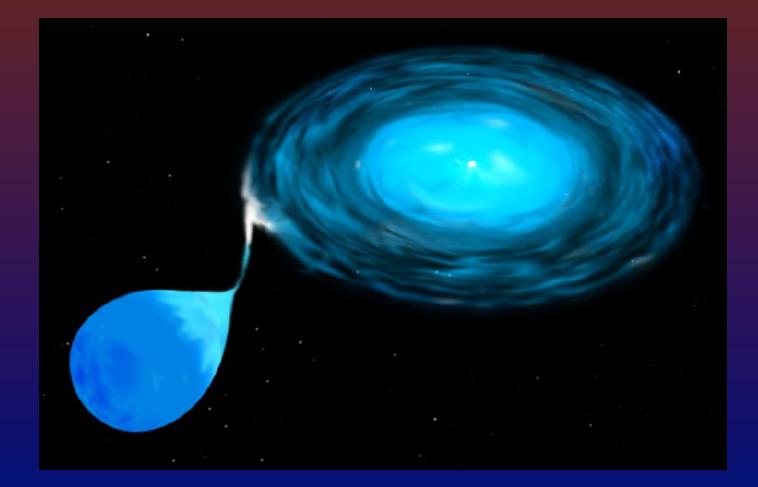
Measurements of Ω and Λ From 42 High-Redshift Supernovae

S. Perlmutter et al.

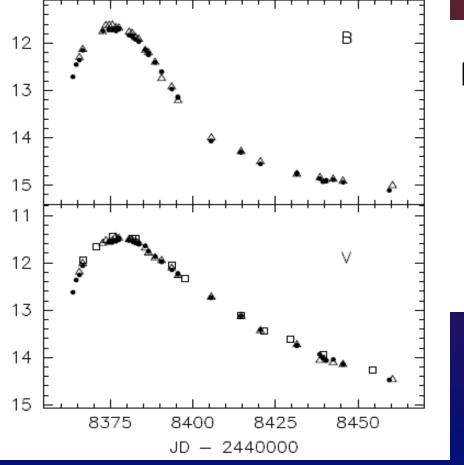
© Anglo-Australian Observatory

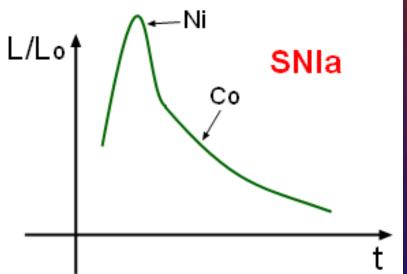


Demise of a Very hungry White Dwarf



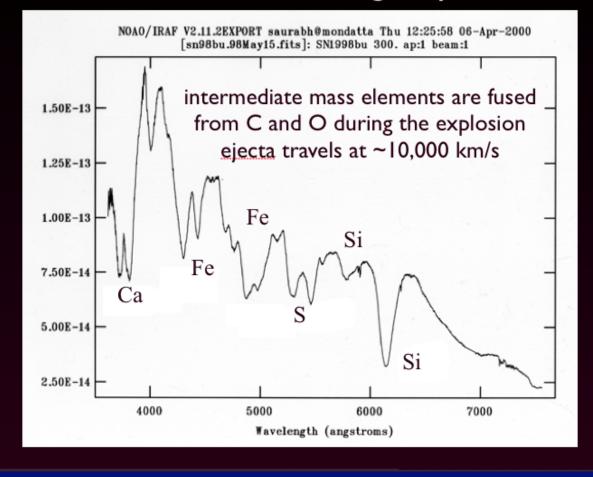
Supernova Light-Curve (Phillips et. al. 1992)





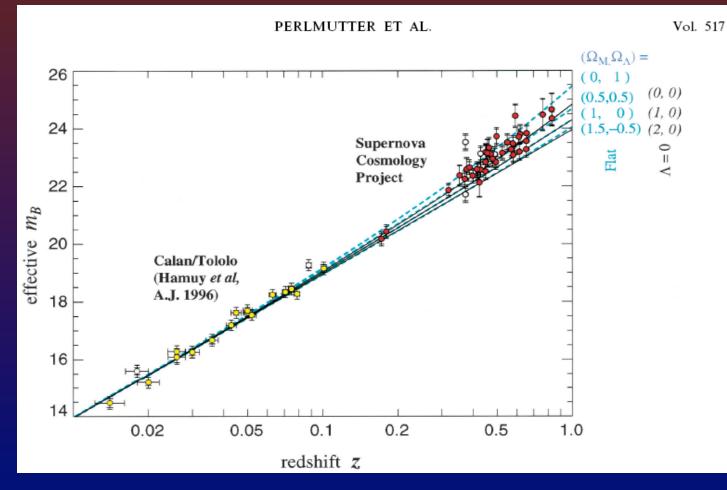
Spectrum of Ia Supernovae

SN 1998bu maximum light spectrum



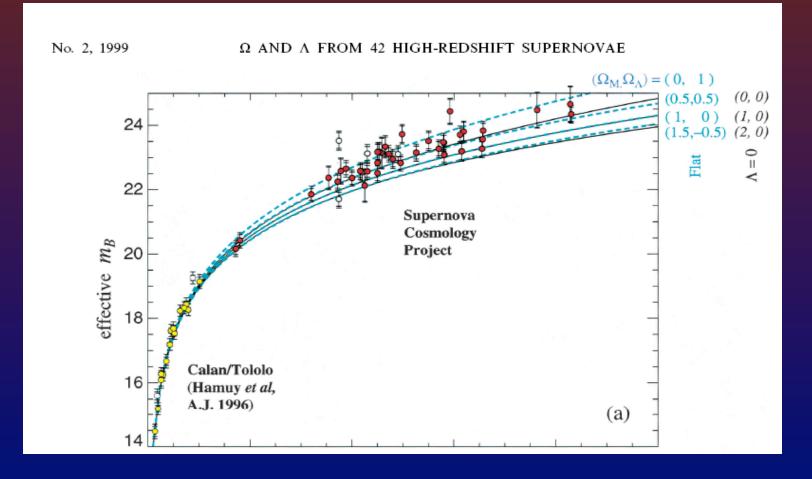


Hubble Diagram For the Ia Supernovae Log Plot

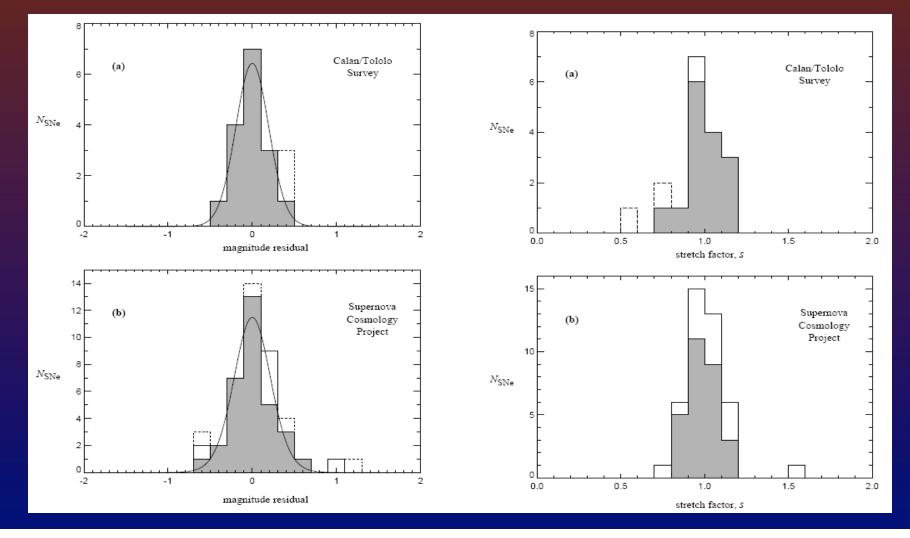




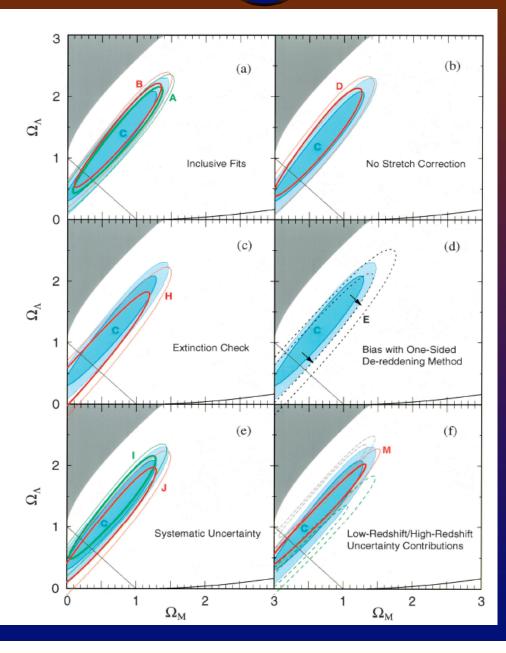
Linear Hubble Diagram For the Ia Supernovae



Checking Uncertainties

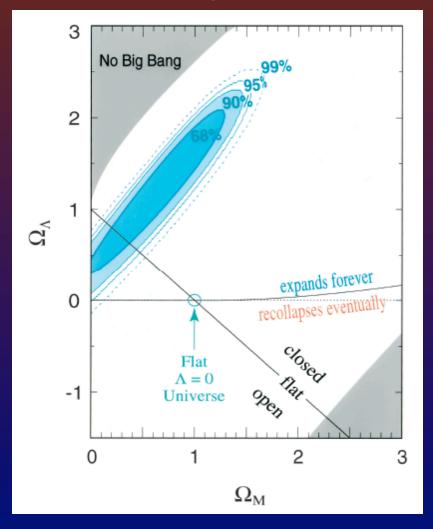


\leq **Best Fit for \Omega** ^{M-}



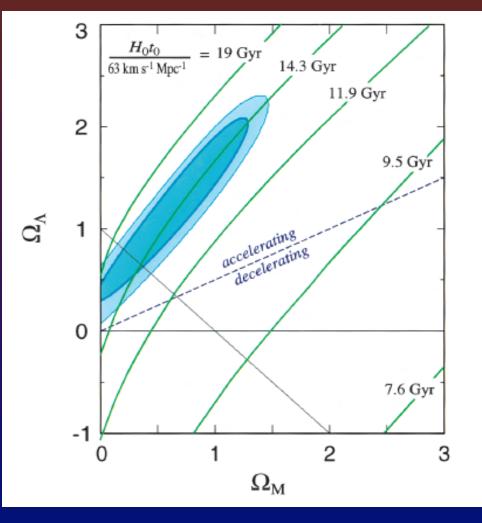


Confidence Regions for Ω_{M} - Ω_{Λ}

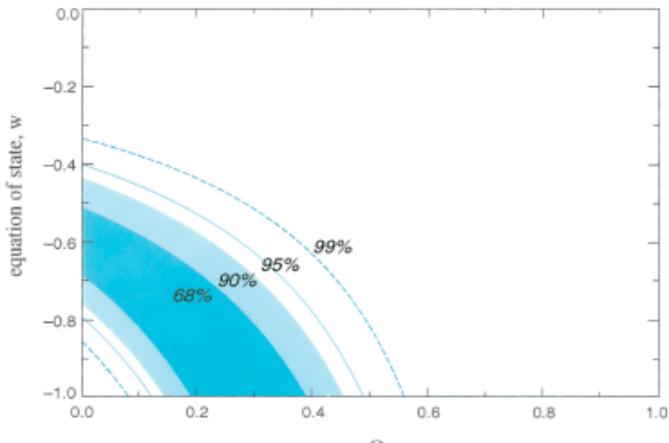




Isochrones of constant Hoto



<u>Time Independent Equation of State</u> <u>for a Flat Space-Time Geometry</u>



 Ω_{M}



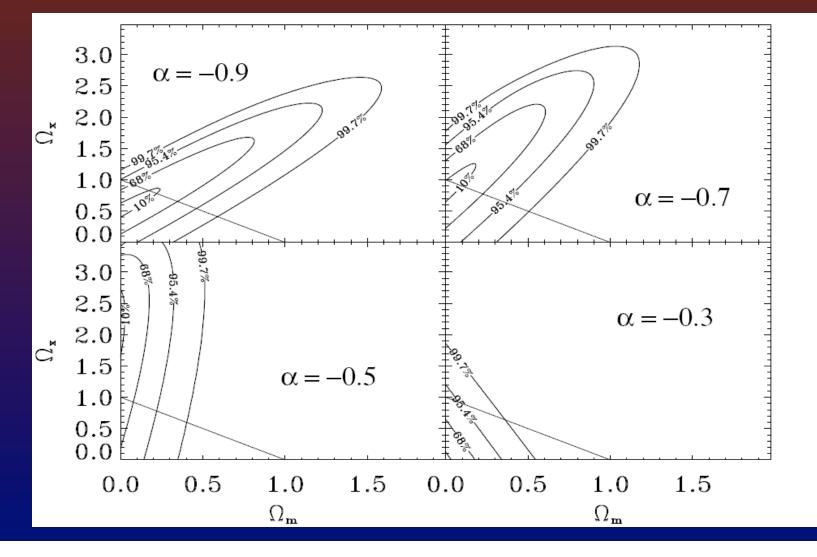
<u>Supernova Limits on the Cosmic</u> <u>Equation of State</u>

Peter M. Garnavich, Saurabh Jha et. al.

<u>Friedmann-Robertson-Walker</u> geometry dependent Luminosity <u>distance</u>

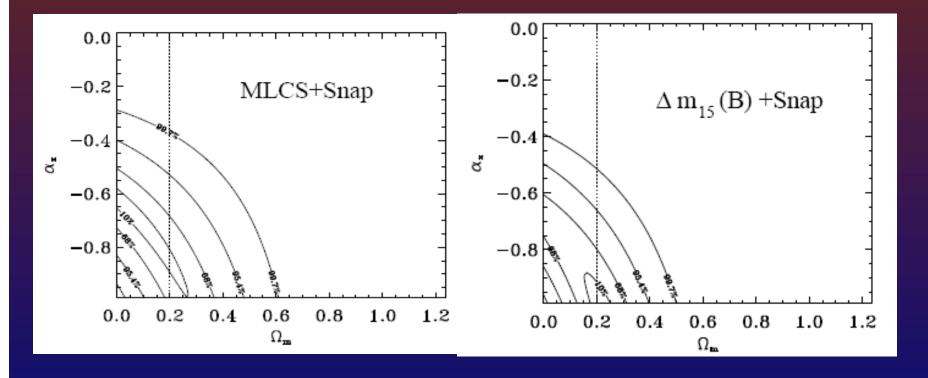
$m = M + 5 \log D_L$	+ 25
$D_L = \frac{c(1+z)}{H_0\sqrt{ \Omega_k }} \operatorname{sinn}$	
$\times \left\{ \sqrt{ \Omega_k } \int_0^z \left[\sum_i \Omega_i (1+z')^{3(1+\alpha_i)} + \Omega_k (1+z')^2 \right]^{-1} \right]^{-1}$	$\left. \frac{1}{2}dz'\right\}$,
$(\sinh(x), \text{ if } \Omega_k > 0;$	(1)
$\operatorname{sinn}(x) = \begin{cases} \sinh(x), & \text{if } \Omega_k > 0 ; \\ x, & \text{if } \Omega_k = 0 ; \\ \sin(x), & \text{if } \Omega_k < 0 , \end{cases}$	(2)
$\sin(x)$, if $\Omega_k < 0$,	

Joint Probability Distributions for Unknown Ω_x Vs. Ω_m at various α 's

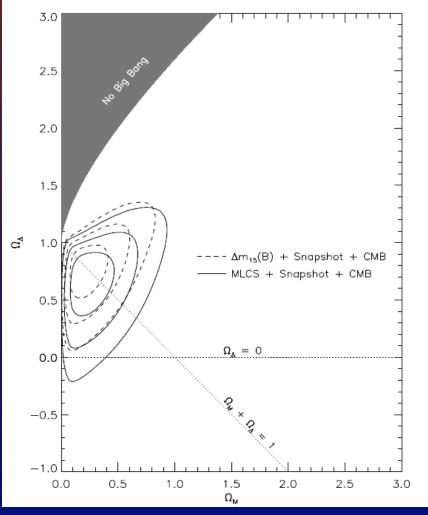




$\frac{\text{Joint Probability Distributions for }\Omega_{\text{m}}}{\text{and Equation of State }\alpha_{\text{x}}}$



Constraints on first Doppler Peak of the CMB Angular Power Spectrum





CMB Confirmation

