

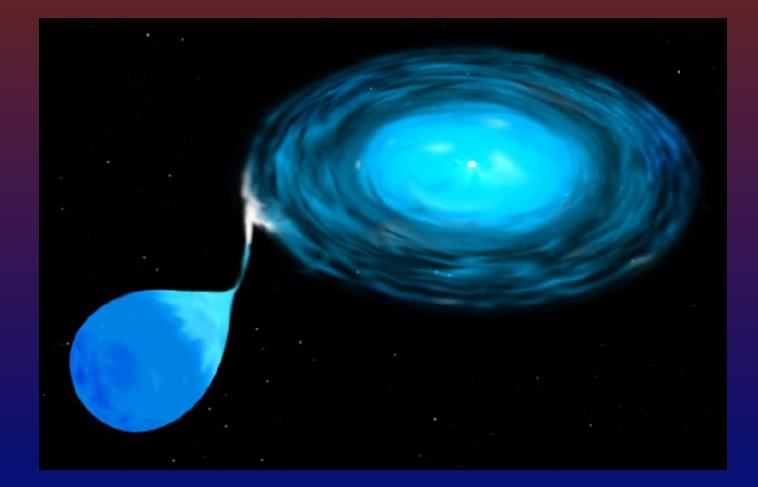
# Measurements of Ω and Λ From 42 High-Redshift Supernovae

S. Perlmutter et al.

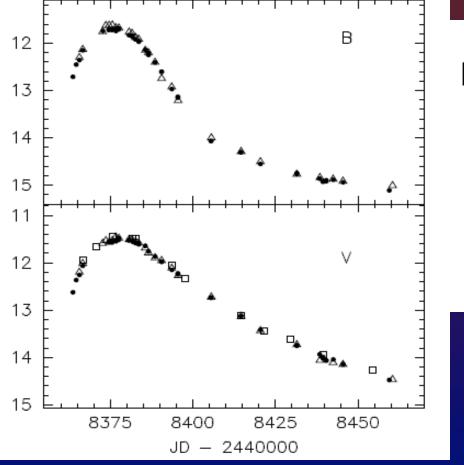
© Anglo-Australian Observatory

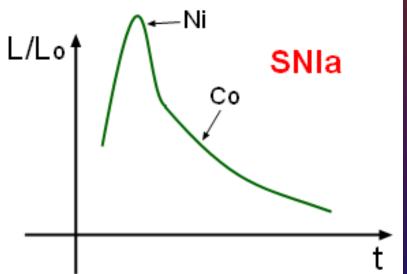


## Demise of a Very hungry White Dwarf



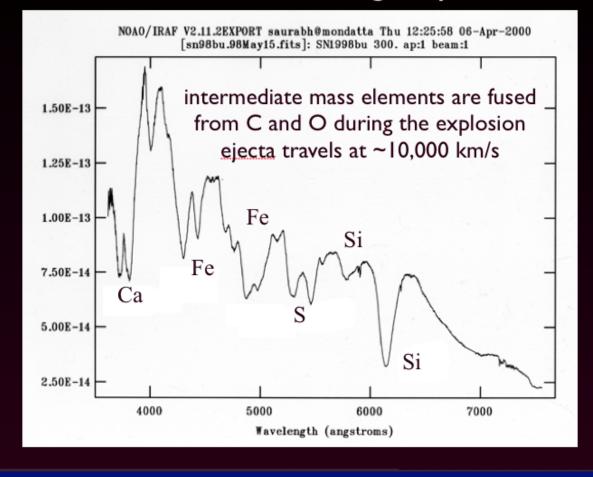
## Supernova Light-Curve (Phillips et. al. 1992)





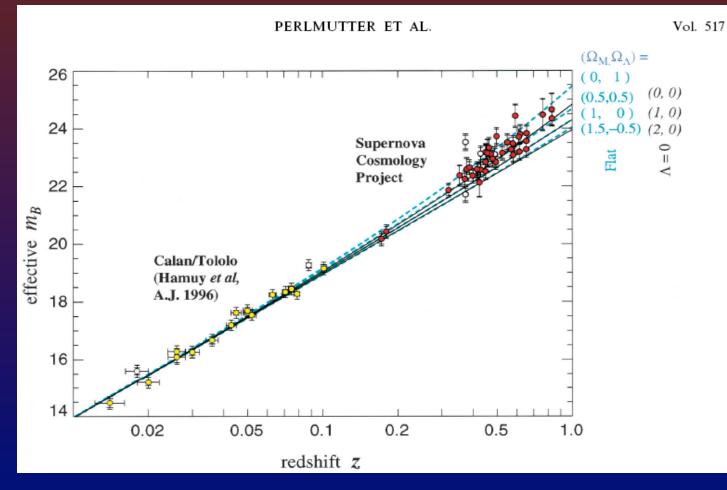
## Spectrum of Ia Supernovae

#### SN 1998bu maximum light spectrum



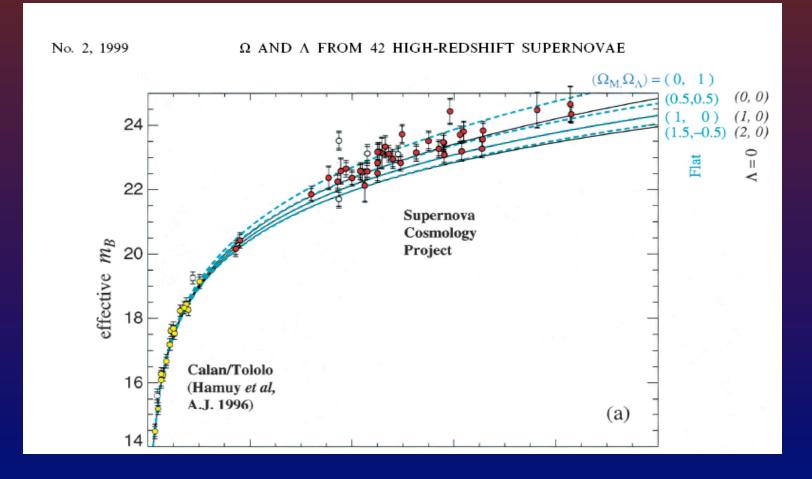


## Hubble Diagram For the Ia Supernovae Log Plot

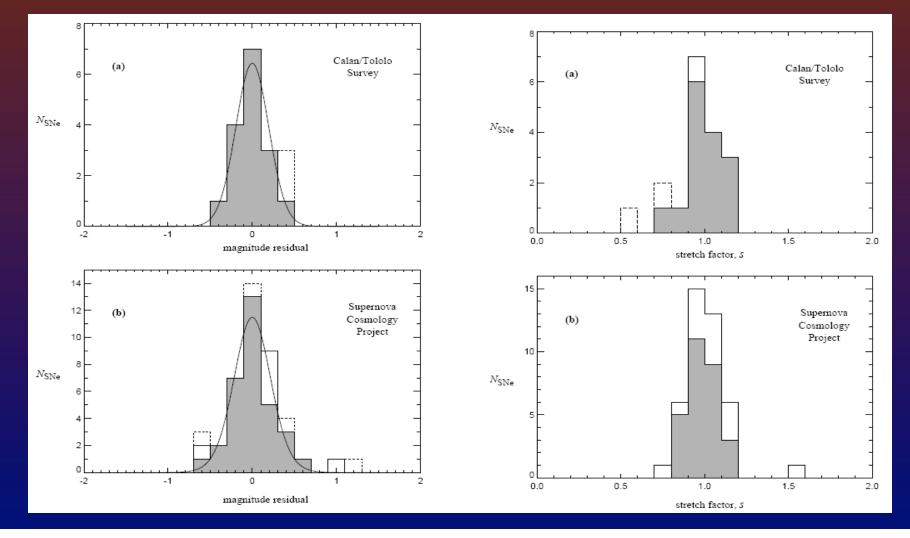




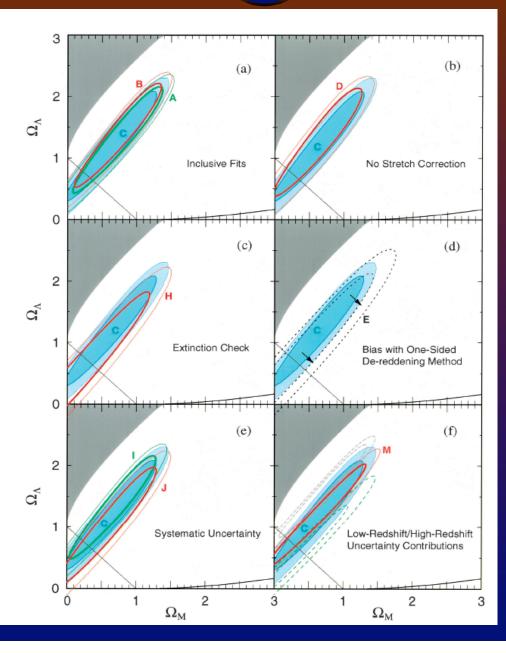
## Linear Hubble Diagram For the Ia Supernovae



### Checking Uncertainties

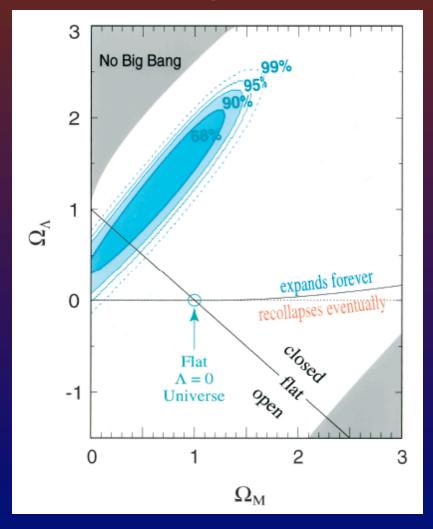


# $\leq$ **Best Fit for \Omega** <sup>M-</sup>



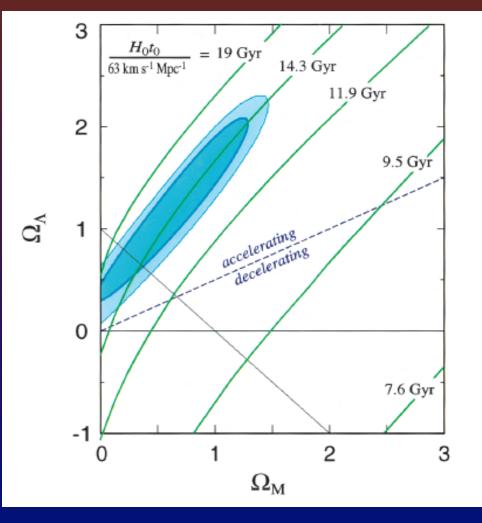


#### Confidence Regions for $\Omega_{M}$ - $\Omega_{\Lambda}$

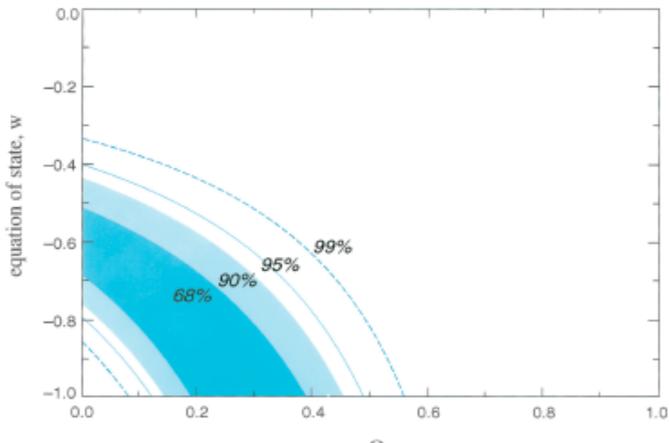




#### Isochrones of constant Hoto



## <u>Time Independent Equation of State</u> <u>for a Flat Space-Time Geometry</u>



 $\Omega_{M}$ 



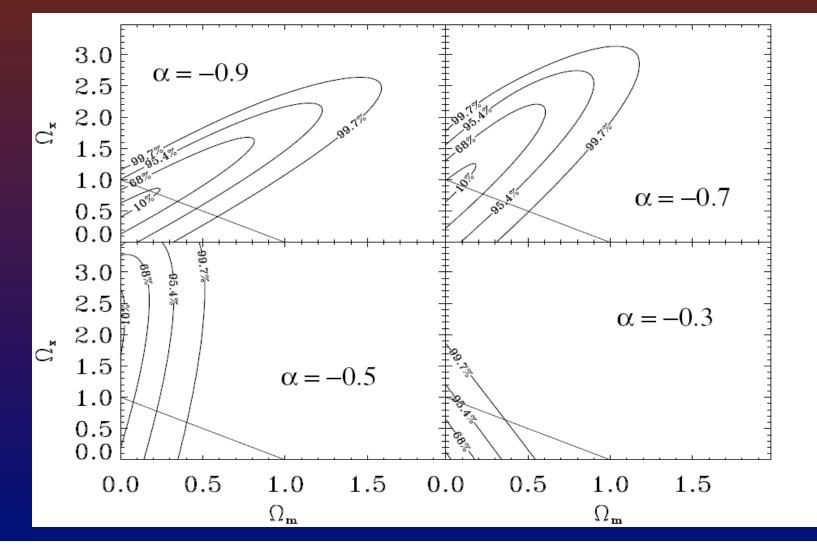
## <u>Supernova Limits on the Cosmic</u> <u>Equation of State</u>

#### Peter M. Garnavich, Saurabh Jha et. al.

## <u>Friedmann-Robertson-Walker</u> geometry dependent Luminosity <u>distance</u>

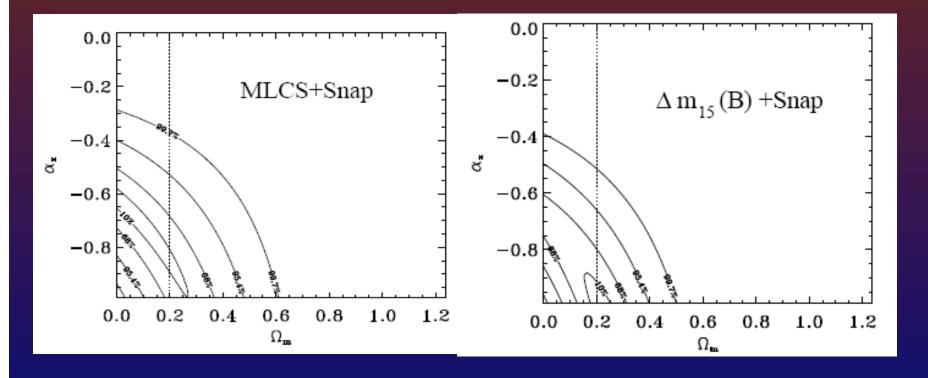
$m = M + 5 \log D_L$	+ 25
$D_L = \frac{c(1+z)}{H_0\sqrt{ \Omega_k }} \operatorname{sinn}$	
$\times \left\{ \sqrt{ \Omega_k } \int_0^z \left[ \sum_i \Omega_i (1+z')^{3(1+\alpha_i)} + \Omega_k (1+z')^2 \right]^{-1} \right]^{-1}$	$\left. \frac{1}{2}dz'\right\}$ ,
$(\sinh(x), \text{ if } \Omega_k > 0;$	(1)
$\operatorname{sinn}(x) = \begin{cases} \sinh(x), & \text{if } \Omega_k > 0 ; \\ x, & \text{if } \Omega_k = 0 ; \\ \sin(x), & \text{if } \Omega_k < 0 , \end{cases}$	(2)
$\sin(x)$ , if $\Omega_k < 0$ ,	

## Joint Probability Distributions for Unknown $\Omega_x$ Vs. $\Omega_m$ at various $\alpha$ 's

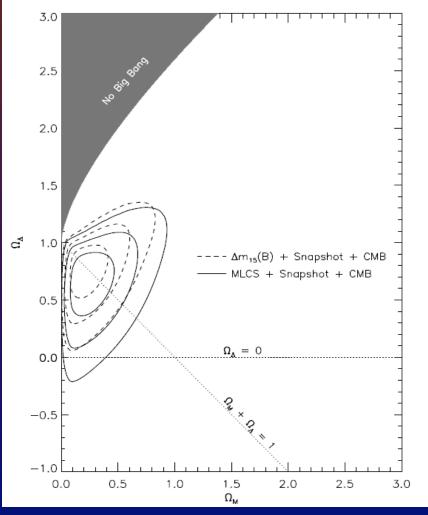




## $\frac{\text{Joint Probability Distributions for }\Omega_{\text{m}}}{\text{and Equation of State }\alpha_{\text{x}}}$



## Constraints on first Doppler Peak of the CMB Angular Power Spectrum





#### **CMB** Confirmation

