Mond



Review of the Sanders & McGaugh 2002, "Modified Newtonian Dynamics as an Alternative to Dark Matter," ARAA, 40, 263 by Dusan Maletic



Phenomenology



Bizarre matter or ordinary field?



Algorithm

Historical issues

 $F=(GM/R^2) f(R/R_o) f(x)=1$ for x<<1 or f(x)=x for x>>1

Tully-Fisher: L~v⁴ larger galaxy/larger discrepancy

Slope of Tully-Fisher: not length but acceleration scale modification



M/L~1/a



Consequences

 $man(a/a_o) = F$ or $gn(|g|/a_o) = g_N$ where n(x) = 1 for x <<1 or n(x) = x for x >>1

- By design: $v^4 = GMa_o$
- Easily falsified by observation.
- $S_m = a_o/G$ surface density
- LSB prediction
- Automatic LSB/HSB distinction
- Freeman's law : brightness limit
- Faber-Jackson : L~ v⁴





Here Lay Dragons

External field effect is required by Mond:

$m(|g_e + g_i|/a_o) g_i = g_{Ni}$

Relation to the Hubble constant

$$a_o = cH_o$$



Galactic scale

Light traces mass (M/L constant for one galaxy) Ho at best known value 75 km/s Mpc



Figure 5 MOND fits to the rotation curves of spiral galaxies with published data, from



Smaller than galactic scales

Pioneer anomaly: $a_p = 8 \ 10^{-8} \ cm/s^2$



 $\log(\sigma_1) (km/s)$





Larger than galactic scale



Mond predicts more mass in clusters Strong lensing observations in clusters However Faber-Jackson still applies



State of theory

Non-relativistic theory: conservation of momentum and energy As well as no dependence of the external field on the inner structure

$$\nabla \cdot \left[\mu \left(\frac{|\nabla \phi|}{a_o} \right) \nabla \phi \right] = 4\pi \, G\rho$$

Relativistic theory: Evolving. After this paper where Scalar-tensor theory is proposed, more advanced TeVeS Treatment have been published. Problems still exist:

Lensing

MOND as modified inertia?



Cosmology

No complete theory, so no real Cosmology.

However, speculative MONDIAN-like Friedmann equation:

$$\dot{r} = u_i^2 - \left[2\Omega_m H_o^2 r_o^3 a_o\right]^{1/2} \ln(r/r_i),$$





Conclusion

Single parameter

Preferred surface density

Pressure supported isotermal systems on vast range of scales

Luminosity-velocity dispersion relation

Problems: not well developed, lensing, large scale

Future: Does MOND represent influence of Cosmology On the local particle dynamics?