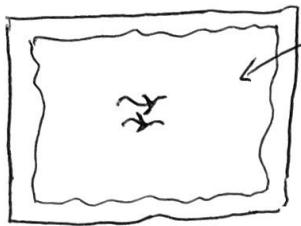


Lecture 9

Black-body radiation

EM radiation enclosed in volume V

@ temp. T : "blackbody cavity"



radiation @ equil. inside
the cavity due to
emission/absorption of
photons by the cavity
walls

Photons of freq. ω :

$$\left\{ \begin{array}{l} E = \hbar \omega, \\ \vec{p} = \hbar \vec{k} \\ \vec{\epsilon} \cdot \vec{k} = 0, \quad |\vec{\epsilon}| = 1 \end{array} \right. \quad |\vec{k}| = \frac{\omega}{c} \quad \Leftrightarrow \vec{E}(\vec{r}, t) = \vec{\epsilon} e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

↑ polarization vector
 \vec{E} field is transverse

use $V = L^3$, periodic BCs:

$$\vec{k} = \frac{2\pi \vec{n}}{L} \Rightarrow \vec{k} \cdot \vec{L} \Rightarrow k_x L = 2\pi \underbrace{n_x}_0, \pm 1, \pm 2, \dots$$

(e.g.
 $(L, 0, 0)$)

Thus, $\vec{n} = (n_x, n_y, n_z)$ where each
vector component $= 0, \pm 1, \pm 2, \dots$

The # of states in ~~observes may really do this~~
is given by:

~~observes may really do this~~
~~observes may really do this~~
~~observes may really do this~~
~~observes may really do this~~ $(n, n+dn)$

$$\cancel{\int_{\text{volume}} \sum_n n^2 d\tau} = \frac{V}{(2\pi)^3} \sum_k k^2 dk$$

Then $E = \sum_{\vec{k}, \vec{\epsilon}} \hbar \omega n_{\vec{k}, \vec{\epsilon}}$ $n_{\vec{k}, \vec{\epsilon}} = 0, 1, 2, \dots$

Next, $Q = \sum_{\{\vec{k}, \vec{\epsilon}\}} e^{-\beta \sum_{\vec{k}, \vec{\epsilon}} \hbar \omega n_{\vec{k}, \vec{\epsilon}}} =$
 $= \prod_{\vec{k}, \vec{\epsilon}} \sum_{n=0}^{\infty} e^{-\beta \hbar \omega n}$ $= \prod_{\vec{k}, \vec{\epsilon}} \frac{1}{1 - e^{-\beta \hbar \omega}}$.
 $\underbrace{\quad}_{\text{geom. series}}$ $\underbrace{\quad}_{\omega = \omega_{\vec{k}}}$

$$\log Q = - \sum_{\vec{k}, \vec{\epsilon}} \log (1 - e^{-\beta \hbar \omega}) = -2 \sum_{\vec{k}} \log (1 - e^{-\beta \hbar \omega}).$$

Furthermore, \downarrow occupation # for photons of momentum \vec{k} , both polarizations

$$\langle n_{\vec{k}} \rangle = \frac{\partial}{\partial(-\beta \hbar \omega)} \log Q = - \frac{2}{1 - e^{-\beta \hbar \omega}} (-e^{-\beta \hbar \omega}) =$$

$$= \frac{2}{e^{\beta \hbar \omega} - 1}.$$

\equiv
Internal energy: $U = \frac{\partial}{\partial(-\beta)} \log Q =$

~~$$= -2 \sum_{\vec{k}} \frac{1}{1 - e^{-\beta \hbar \omega}} (-\hbar \omega) =$$~~

$$= \sum_{\vec{k}} \hbar \omega \langle n_{\vec{k}} \rangle.$$

Next, express

$$\log Q = -2 \sum_{\vec{k}} \log \left(1 - e^{-\beta \hbar c \frac{\underbrace{2\pi |\vec{n}|}_{\text{in}}}{\sqrt{3}}} \right).$$

Then $p = \frac{1}{\beta} \frac{\partial \log Q}{\partial V} = -\frac{2}{\beta} \sum_{\vec{k}} \frac{1}{1 - e^{-\beta \hbar \omega}} (-e^{-\beta \hbar \omega}) \times$

$$\times (-\beta \hbar c \times 2\pi |\vec{n}|) \left(-\frac{1}{3\sqrt{3}} \right) = \frac{1}{|\vec{k}|}$$

$$= \sum_{\vec{k}} \langle n_{\vec{k}} \rangle \frac{\hbar c \times \left(\frac{2\pi |\vec{n}|}{\sqrt{3}} \right)}{3V} = \frac{1}{3V} \underbrace{\sum_{\vec{k}} \hbar \omega \langle n_{\vec{k}} \rangle}_{U}, \text{ s.t.}$$

$$pV = \frac{U}{3}$$

Now, consider $(V \rightarrow \infty)$

$$U = \frac{2V}{(2\pi)^3} \int_0^{\infty} dk (4\pi k^2) \frac{\hbar ck}{e^{\beta \hbar ck} - 1} =$$

$$= \frac{V \hbar}{\pi^2 c^3} \int_0^{\infty} d\omega \frac{\omega^3}{e^{\beta \hbar \omega} - 1}.$$

In other words,

$$\frac{U}{V} = \int_0^{\infty} d\omega U(\omega, T)$$

$$\frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1} \quad \begin{matrix} \leftarrow \\ \text{Planck's radiation} \\ \text{law} \end{matrix}$$

$\underbrace{\frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}}$ energy density of photons of freq. ω , regardless of polarization & direction of \vec{k}

The $\int_0^\infty d\omega u(\omega, T)$ can be evaluated,

$$\text{yielding } \frac{u}{V} = \frac{\pi^2}{15} \frac{(k_B T)^4}{(\hbar c)^3} \sim T^4$$

$\frac{u}{V} \rightarrow \infty$ as $T \rightarrow \infty$ since the number of photons is not bounded from above.

Now, consider radiation from a small opening in the wall of the black-body cavity:



Energy radiated per second via photons of frequency ω :

$\frac{1}{2}$ of the photons point away from A



$$I(\omega, T) = \frac{c}{2} A u(\omega, T) \underbrace{\frac{1}{4\pi} \int d\Omega \cos\theta}_{2\pi} = \frac{c}{4} A u(\omega, T)$$

two hemispheres, unit freq. \rightarrow unit solid angle

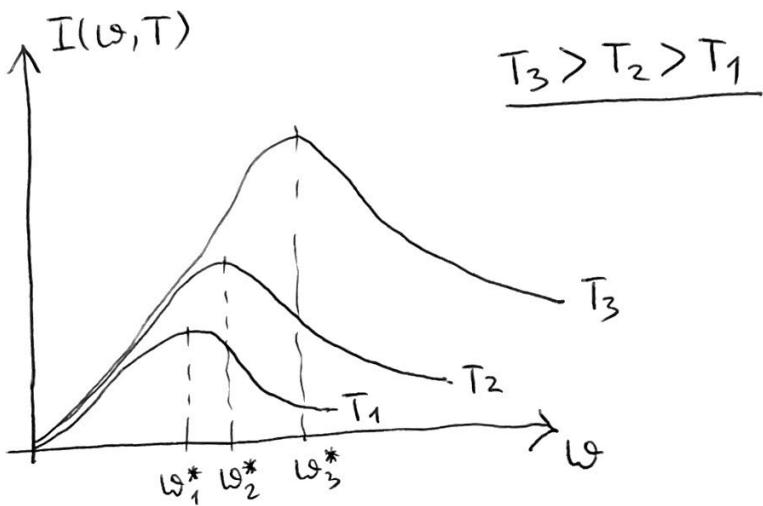
$$\int d\Omega \cos\theta = 2\pi \int_0^{\pi/2} d\theta \times \sin\theta \cos\theta = 4\pi \int_0^{\pi/2} d(-\cos\theta) \cos\theta =$$

$$= 4\pi \int_0^1 du u = 2\pi$$

Hence

$$I(T) = \int_0^\infty d\omega \frac{I(\omega, T)}{A} = \frac{C}{4} \int_0^\infty d\omega u(\omega, T) \quad \text{=} \\ \underbrace{\text{flux}}_{\text{per unit}} \underbrace{\text{area of the opening}}_{\frac{\pi^2}{15} \frac{(k_B T)^4}{(\hbar c)^3}}$$

$$\text{=} \underbrace{\frac{\pi^2 k_B^4}{60 \hbar^3 c^2} T^4}_{6, \text{ Stefan's constant}} \Rightarrow [I(T) = 6 T^4]$$



Area under each ~~curve~~ curve
is $6T_i^4$ ($i=1,2,3$)

Note Consider

$$\frac{u}{V} = \frac{\hbar}{\pi^2 c^3} \int_0^\infty d\omega \frac{\omega^3}{e^{\beta \hbar \omega} - 1} = \left(\frac{\hbar}{\pi^2 c^3} \right) \left(\frac{1}{\beta \hbar} \right)^4 \times$$

$$\times \int_0^\infty du \frac{u^3}{e^u - 1} = G \frac{(k_B T)^4}{(\hbar c)^3} \frac{1}{\pi^2} = \frac{\pi^2}{15} \frac{(k_B T)^4}{(\hbar c)^3} \quad \text{as before}$$

$\underbrace{G \left[= \frac{\pi^4}{15} \right]}_{\text{}} \quad \text{=}$

Phonons in Solids

Phonons = quanta of sound waves in a macroscopic solid.

Hamiltonian \approx sum of harmonic oscillators corresponding to normal modes

Quantization: $\tilde{E} e^{i(\vec{k} \cdot \vec{r} - \omega t)} \Leftrightarrow$ phonons, the ω_j
 $\omega = |\vec{k}| c$
↑ speed of sound

Polarization \tilde{E} : 2 transverse + 1 longitudinal mode

Solid with N atoms $\Rightarrow 3N$ normal modes:

$$\omega_1, \omega_2, \dots, \omega_{3N}$$

Each freq. ω_j can characterize multiple phonons; the total # of phonons is not conserved.

We focus on the Debye model: normal modes are the lowest $3N$ frequencies in an elastic solid of volume V :
periodic BCs $\Rightarrow \vec{k} = \frac{2\pi}{L} \vec{n}$, $L = V^{1/3}$
and $\vec{n} = (n_1, n_2, n_3)$, with $n_i = 0, \pm 1, \pm 2, \dots$
 $i = 1, 2, 3$

Then # normal modes in the $(\omega, \omega + d\omega)$ interval:

$$f(\omega) d\omega = \frac{3V}{(2\pi)^3} 4\pi k^2 dk = \sqrt{\frac{3\omega^2}{2\pi^2 C^3}} d\omega.$$

$k = \frac{\omega}{C}$

$\omega_m \leftarrow \text{max freq.}$

Next, $\int_0^{\omega_m} d\omega f(\omega) = 3N$, or

$$\frac{3V}{2\pi^2 C^3} \int_0^{\omega_m} d\omega \omega^2 = 3N,$$

$\underbrace{\omega_m^3}_{\frac{1}{3}}$

$$\omega_m^3 = \frac{6\pi^2 C^3}{N}, \quad N = \frac{V}{\lambda^3}.$$

Correspondingly,

$$\lambda_m = \frac{2\pi C}{\omega_m} = \left(\frac{4}{3}\pi N\right)^{1/3} \sim \underbrace{N^{1/3}}_{\substack{\uparrow \\ \text{min wavelength}}} \quad \text{interparticle distance}$$

Total energy :

$$E\{n_i\} = \sum_{i=1}^{3N} n_i \hbar \omega_i$$

$$Q = \sum_{\{n_i\}} e^{-\beta E\{n_i\}} = \prod_{i=1}^{3N} \frac{1}{1 - e^{-\beta \hbar \omega_i}}, \text{ or}$$

$$\log Q = - \sum_{i=1}^{3N} \log(1 - e^{-\beta \hbar \omega_i})$$

Furthermore,

$$\langle n_i \rangle = \frac{\partial}{\partial(-\beta\hbar\omega_i)} \log Q = \frac{1}{e^{\beta\hbar\omega_i} - 1}.$$

Internal energy:

$$U = -\frac{\partial}{\partial\beta} \log Q = \sum_{i=1}^{3N} \hbar\omega_i \langle n_i \rangle = \\ = \sum_{i=1}^{3N} \frac{\hbar\omega_i}{e^{\beta\hbar\omega_i} - 1}$$

In the continuous limit, $(v \rightarrow \infty)$

$$U = \frac{3V}{2\pi^2 C^3} \int_0^{\omega_m} d\omega \omega^2 \frac{\hbar\omega}{e^{\beta\hbar\omega} - 1}, \text{ or} \\ U = \underbrace{\frac{3V}{2\pi^2 C^3} \frac{(k_B T)^4}{\hbar^3}}_{\text{constant}} \int_0^{\beta\hbar\omega_m} du \frac{u^3}{e^u - 1}.$$

Note that

$$\frac{9N(k_B T)^4}{(\hbar\omega_m)^3} = \frac{9N(k_B T)^4}{\hbar^3 \frac{6\pi^2 C^3}{V}} = \frac{3V}{2} \frac{1}{\pi^2 C^3} \frac{(k_B T)^4}{\hbar^3}.$$

Thus, $\underbrace{\frac{u}{N}}_{\text{per atom}} = \frac{9(k_B T)^4}{(\hbar\omega_m)^3} \int_0^{\beta\hbar\omega_m} du \frac{u^3}{e^u - 1}.$

Introduce the Debye function

$$D(x) = \frac{3}{x^3} \int_0^x dt \frac{t^3}{e^t - 1} = \begin{cases} 1 - \frac{3}{8}x + \frac{1}{20}x^2 + \dots & x \ll 1 \\ \frac{\pi^4}{5x^3} + O(e^{-x}) & x \gg 1 \end{cases}$$

and the Debye T:

$$k_B T_D \equiv \hbar \omega_m = \hbar c \left(\frac{6\pi^2}{v} \right)^{1/3} \Rightarrow \beta \hbar \omega_m = \frac{T_D}{T},$$

$$\text{Then } \frac{u}{N} = 3(k_B T) \underbrace{\frac{3}{(T_D/T)^3} \int_0^{T_D/T} du \frac{u^3}{e^u - 1}}_{D(\frac{T_D}{T})} \quad \textcircled{=}$$

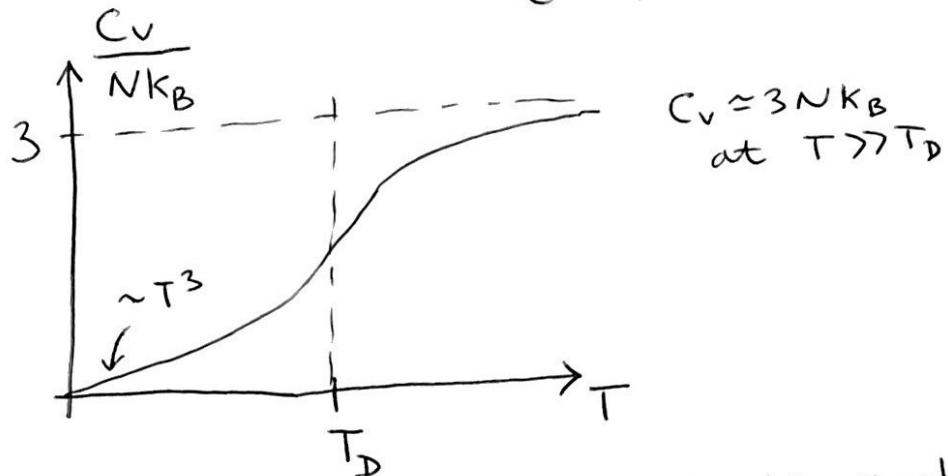
$$\textcircled{=} 3(k_B T) D\left(\frac{T_D}{T}\right) \begin{cases} 3k_B T \left(1 - \frac{3}{8} \frac{T_D}{T} + \dots\right) & T \gg T_D \\ 3k_B T \frac{\pi^4}{5} \left(\frac{T}{T_D}\right)^3 & T \ll T_D \end{cases}$$

The specific heat:

$$\begin{aligned} \frac{C_V}{Nk_B} &= 3D\left(\frac{T_D}{T}\right) + 3T \frac{dD\left(\frac{T_D}{T}\right)}{dT} = \\ &= 3D\left(\frac{T_D}{T}\right) + 3T \frac{\frac{dD\left(\frac{T_D}{T}\right)}{d\left(\frac{T_D}{T}\right)} \frac{dT}{d\left(\frac{T_D}{T}\right)}}{= \\ &= 3D\left(\frac{x}{T}\right) + 3T \frac{d}{dx} \left[\frac{3}{x^3} \int_0^x du \frac{u^3}{e^u - 1} \right] \times \left(-\frac{T_D}{T^2}\right) = \\ &= 3D(x) - 3x \left[\frac{3(-3)}{x^4} \int_0^x du \frac{u^3}{e^u - 1} + \frac{3}{x^3} \frac{x^3}{e^x - 1} \right] = \\ &= 3D(x) + 9D(x) - \frac{9x}{e^x - 1} = 3 \left(4D(x) - \frac{3x}{e^x - 1} \right). \end{aligned}$$

This leads to

$$\frac{C_V}{Nk_B} = \begin{cases} 3 \left[1 - \frac{1}{20} \left(\frac{T_D}{T} \right)^2 + \dots \right] & (x \ll 1) \\ \underbrace{\frac{12\pi^4}{5} \left(\frac{T}{T_D} \right)^3}_{\text{here, } \frac{x}{e^{x-1}} \rightarrow 0 \text{ exponentially}} & T \gg T_D \\ & T \ll T_D \\ & (x \gg 1) \end{cases}$$



For many solids, $T_D \approx 200\text{K}$ and $C_V \approx 3Nk_B$ @ $T \approx 300\text{K}$